Bulk boundary layer concepts for simplified models of tropical dynamics

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Abstract We review bulk representations of tropical and subtropical maritime atmospheric boundary layers. Three types of bulk representations are studied in detail: stratocumulus topped mixed layers; trade-wind layers; and sub-cloud mixed layers. Through the development of a consistent description of these disparate regimes, connections among their varied representations are emphasized, as well as their relation to regions of deeper convection. New results relating to the equilibrium mass flux and cloud fraction in the trade-winds; the ability of bulk models to qualitatively represent major cloud regimes; and the relationship amongst different bulk representations of the surface layer are presented. Throughout we emphasize the identification of consistent and physically based mixing and cloud regime rules for use in intermediate complexity models of the tropical climate, which in turn can be used to study cloud and dynamical interactions on larger scales.

Keywords Atmospheric Boundary Layer · Entrainment · Cumulus Mass Flux

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1 Introduction

Bulk, or integral, models of the atmospheric boundary layer (ABL) have long proven useful in a variety of context. Indeed the very definition of the boundary layer depth in classical boundary layer theory is in terms of an integral over the depth of a layer [34]. Bulk ABL models have appeared as a basis for parameterizing the ABL in larger-scale models [11,3], and as essential elements in simple thermodynamic models of the tropical climate system[7,24]. They are often also used as the basis for diagnostic studies of the ABL [1] and particular processes, or ABL regimes, such as stratocumulus [16, 18,47], surface winds over the tropical ocean [43,23], trade-wind cloudiness [5,6,27], or deep convection [13,32]. A seminal study of this framework is one component of the landmark paper by Arakawa and Schubert [3]. Since this study, an ever increasing variety of bulk models has been developed, and our physical insight into the basic processes parameterized in these models has deepened, which largely motivates our current investigation.

In this paper we draw on this rich literature of bulk, or integral, representation of the ABL and attempt to extract those ideas of most relevance to the representation of tropical circulations. Our idea is to provide a unified framework through which varied, and sometimes, quite specialized ideas can be described, and relationships amongst seemingly disparate lines of thought can be clarified.

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The hope is that a unified presentation of the mathematical structure of past ideas, and a careful consideration of important (and in some cases novel) limits will make this body of thought both more understandable, and accessible, to the broader scientific community. Although partly a review, we do not strive to be comprehensive. Instead, we focus on the development of the theory, particularly in so far as the introduction of simplifying assumptions or new physical insights can help clarify its basic structure, and by implication that of the physical system; also in this respect a number of new results are presented.

The emphasis of this study is on developing the theory of bulk representation of the thermodynamics structure of the ABL away from regions of deep convection. Such an emphasis is motivated by the fact that cloud processes in such regions are thought to be critical for climate, and that thermodynamically these regions tend to be most distinct from the free-troposphere, and thus more susceptible to neglect. In §2 we develop the notation used throughout the remainder of the paper. The structure of a generic bulk description of the ABL and its energetics is provided in §3. Section 4 explores the behavior of the bulk theory under constraints appropriate for regions of stratocumulus. In §5 two different approaches for representing trade-cumulus are presented. Ways of using the ideals developed in §4 and §5, coupling them to one another, as well as to boundary layer dynamical processes, and neighboring regions of deep convection are discussed in §6. A brief summary and conclusions are presented in §7.

2 Setup and notation

Throughout we consider the structure of shallow (as compared to an atmospheric scale height) layers, for which the atmospheric analog to the Boussinesq system (wherein the basic state potential temperature Θ rather than actual temperature is constant) provides a good analog to the real fluid. Thus the independent variables are the three components of physical space $\{x,y,z\}$ and time t. A less restrictive, but mathematically equivalent, system of equations follows by working in mass coordinates, in which case pressure p replaces height z as an independent variable, and mass fluxes replace volume fluxes. Our terminology often anticipates this equivalence by equating volume fluxes with mass fluxes—in the Boussinesq system these are linearly related by the basic state density, which is in any case near unity.

We choose to describe the thermodynamic state of our system in terms of the liquid water static energy, $s \equiv c_p T + gz - Lq_l$, the total-water specific humidity q, and the pressure p. Here q_l is the liquid water specific humidity, c_p is the isobaric specific heat, L is the enthalpy of evaporation, and g is the gravitational acceleration. Given the surface pressure, p_0 , the pressure at some height z is given by the hydrostatic constraint, i.e., $p(z) = p_0 + \int_0^z \rho g \, dz$, where the density, ρ , is a function of the state of the system. To describe the energetics of the system buoyancy perturbations must be characterized; for moist Boussinesq systems these are often represented by perturbations in the virtual potential temperature, θ_v , as

$$b' \equiv -g \frac{\rho'}{\rho_{00}} \approx g \frac{T'_v}{T_{00}} \approx g \frac{\theta'_v}{\Theta} \quad \text{with} \quad T_v \equiv \frac{p}{\rho R_d} = T \left[1 + 0.608q - 1.608q_l \right].$$
 (1)

The numerical factors account for the difference of gas constants for vapor and air respectively¹, which arise because the effective gas constant of the ideal fluid depends on the mass fraction of water. Subscript 00 denotes a basic state value. Given the characterization of the state of the system (i.e., s, q and p) in equilibrium all other thermodynamic variables (such as θ_v) are straightforward to calculate. The dynamic state of the system is given by the velocity vector (u, v, w). Because one of the motivations for developing a bulk description of the ABL is to integrate out vertical dependencies, we distinguish between vertical and horizontal motion. Thus \mathbf{u} is reserved for the horizontal velocity, and ∇ is defined to be a two dimensional operator in the horizontal plane, i.e., $\nabla \equiv \{\partial_x, \partial_y\}$. With this notation the Bousinessq continuity equation becomes:

$$\nabla \cdot \mathbf{u} + \partial_z w = 0. \tag{2}$$

Both s and q are chosen as state variables because they are invariant under reversible² displacements of a fluid admitting phase changes. To aid the development of equation-sets which apply equally to

 $^{^{-1}}$ $(R_v - R_d)/R_d \approx 0.608$, with $R_v = 461.5$ J kg⁻¹ and $R_d = 287$ J kg⁻¹ the gas constants for water vapor and dry air respectively.

 $^{^{2}\,}$ Formally s is only conserved is such displacements are also hydrostatic.

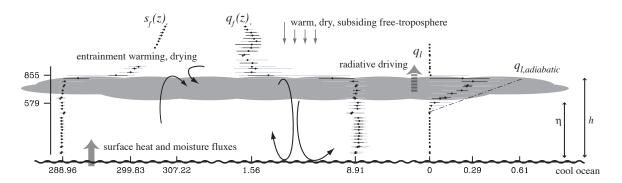


Fig. 1 Structure of the stratocumulus topped boundary layer as observed during DYCOMS-II.

both s and q, (or for that matter any adiabatic, scalar, invariant of the flow) we introduce the symbol ϕ to denote a generic adiabatic invariant, i.e., for the most part $\phi \in \{s, q\}$. An ensemble average is denoted by an over-bar, which defines the deviations (or fluctuations) as follows:

$$\phi' \equiv \phi - \overline{\phi}. \tag{3}$$

The depth of the ABL is denoted by h, and the depth of the sub-cloud layer is measured by η sometimes also called the lifting condensation level. Because we idealize the fluid in a manner which admits discontinuities in state variables at h we develop a notation which admits one-sided limits, such that $h_+ \equiv h + \epsilon$ and $h_- \equiv h - \epsilon$, where ϵ is a vanishingly small (with respect to h) constant. Similarly, subscript "+" or "-" denotes ensemble average values valid at $z = h_+$, or $z = h_-$ respectively. For instance, $\phi_+ = \overline{\phi}(x, y, t; z = h_+)$. Bulk averages are denoted by an over-hat, such that

$$\widehat{\phi}(x,y,t) \equiv \frac{1}{h_{+}} \int_{0}^{h_{+}} \overline{\phi}(x,y,z,t) \, \mathrm{d}z. \tag{4}$$

Note that for this case, as for the subscript "+" or "-" the reference to ensemble averages is implicit. Differences between $\widehat{\phi}$ and values just above h or at the surface are denoted as

$$\Delta_{+}\phi \equiv \phi_{+} - \widehat{\phi} \quad \text{and} \quad \Delta_{0}\phi \equiv \widehat{\phi} - \phi_{0}$$
(5)

respectively, and we further define

$$\Delta \phi \equiv \Delta_{+} \phi + \Delta_{0} \phi. \tag{6}$$

That the Δ symbol operates on the ensemble averaged value of the field is again implicit. In addition, the subscripts f, m, 0 and ∞ are used at times to denote values representative of the free-troposphere, the sub-cloud mixed layer, the surface and in equilibrium respectively. Depending on the context ϕ_f may or may not equal ϕ_+ , likewise ϕ_m may or may not equal $\widehat{\phi}$. Note that in the case when h is associated cloud top $\Delta_+\phi=\phi_f-\widehat{\phi}$, which is not true in cases when h is identified with the height of the sub-cloud layer.

To help summarize some of this notation, Fig. 1 shows the structure of a stratocumulus topped ABL for which $\hat{s}/c_p \approx 289$ K, and s_f varies significantly with height, such that $s_+/c_p \approx 300$ K. $\hat{q} \approx 8.9$ g kg⁻¹ and q_f is approximately constant so that $q_+ \approx 1.6$ g kg⁻¹. s_0 and q_0 are not indicated but correspond to a sea-surface temperature roughly 1-1.5 K warmer than the overlying air. Cloud top is at h=855 m and $\eta \approx 580$ m.

Lastly, in anticipation of the development of equation sets valid for the bulk fluid, we define here the substantial derivative with respect to the bulk average of the ensemble averaged horizontal flow, such that

$$\frac{D}{Dt} \equiv \partial_t + \widehat{\mathbf{u}} \cdot \nabla, \tag{7}$$

where for emphasis it is worth repeating that both \mathbf{u} and ∇ are defined over the horizontal plane.

3 Bulk theory

3.1 Thermodynamic evolution of a variable depth atmospheric boundary layer

Focusing for a moment on the state variables, a traditional Reynolds decomposition into expected values and fluctuations results in the following equation for the expected values,

$$\partial_t \overline{\phi} + \overline{\mathbf{u}} \cdot \nabla \overline{\phi} + w \partial_z \phi = -\nabla \cdot \overline{\mathbf{u}' \phi'} - \partial_z \overline{w' \phi'} - \partial_z \overline{F}_{\phi}, \tag{8}$$

The first two terms on the rhs describe a turbulent flux, which because of molecular processes acts irreversibly from the perspective of the ensemble averaged quantities. The last term on the rhs, involves F_{ϕ} , which denotes a diabatic flux of ϕ not associated with fluid motions, i.e., precipitation or radiative transfer, but not mixing. Its divergence acts as a source of ϕ . To the extent \overline{F}_{ϕ} depends non-linearly on the fluid state it will depend on the fluctuating component of the fluid. Equation (8) is the starting point for almost any analysis of the thermodynamics of high Reynolds number fluids.

In the case when: (i) the turbulent flux of ϕ is horizontally homogeneous; and (ii) baroclinic circulations within the ABL contribute negligibly to the evolution of $\widehat{\phi}$ (i.e., $\widehat{\mathbf{u}\phi} = \widehat{\mathbf{u}}\widehat{\phi}$) the bulk average of Eq. 8 is simply

$$h\frac{D\widehat{\phi}}{Dt} - \Delta_{+}\phi \left[\frac{Dh}{Dt} + \mathcal{D}h \right] = -\Delta \overline{w'\phi'} - \Delta F_{\phi}$$
(9)

where $\mathcal{D} \equiv \nabla \cdot \hat{\mathbf{u}}$ denotes the divergence of the bulk wind. Integrating the continuity equation (2) over the ABL yields $\mathcal{D}h = \overline{w}_+ - \Delta_+ \mathbf{u} \cdot \nabla h$. So $\mathcal{D}h$ is very nearly equal to the value of the vertical velocity at h_+ . Note that in deriving (9) and defining \mathcal{D} it has not been necessary to assume that $\nabla \cdot \overline{\mathbf{u}}$ is constant over the layer as is sometimes thought.

For reasons that will become clear momentarily we define three velocity scales, V, M, and E, to satisfy the following

$$V \equiv -\frac{\overline{w'\phi'}_0}{\Delta_0\phi}, \quad M \equiv -\frac{\overline{w'\phi'}_+}{\Delta_+\phi} \quad \text{and} \quad E = \frac{Dh}{Dt} + \mathcal{D}h + M.$$
 (10)

So doing assumes that the $\overline{w'\phi'}_0$ vanishes with $\Delta_0\phi$ and $\overline{w'\phi'}_+$ vanishes with $\Delta_+\phi$. If we further assume that V and M are the same irrespective of whether ϕ is associated with s or q, then conservation of mass (volume), enthalpy (as measured by \hat{s}) and moisture within the bulk layer follow directly from (9) and (10):

$$\frac{Dh}{Dt} = E - \mathcal{D}h - M \tag{11}$$

$$\frac{D\widehat{s}}{Dt} = \frac{E\Delta_{+}s - V\Delta_{0}s - \Delta F_{s}}{h} \tag{12}$$

$$\frac{D\widehat{q}}{Dt} = \frac{E\Delta_{+}q - V\Delta_{0}q - \Delta F_{q}}{h} \tag{13}$$

These three equations are the basis for all consistent integral treatments of the thermodynamic state of the ABL. As written above they involve 14 unknowns so 11 auxiliary statements are required for their closure. By specifying the state of the large-scale atmosphere and ocean (effectively the boundary conditions) one in effect specifies $\hat{\mathbf{u}}, s_0, q_0, s_+, q_+$ and closure only requires models of F_s, F_q, V, M , and E. Typically F_q is associated with precipitation, while F_s will depend on the combined effects of radiation and precipitation. Both processes act non-locally by transporting enthalpy across fluid streamlines. In the subsequent exposition we will mostly consider the non-precipitating limit, i.e., $\overline{F}_q = 0$, in which case \overline{F}_s only represents the radiative flux divergence across the layer, a process which is relatively well described by our present understanding of radiative transfer given knowledge of the layer's state. The exchange velocities, V, M and E require more discussion and are treated in turn below.

3.1.1 Surface exchange, V

From Eq. 10, V defines the surface fluxes. Of the three velocity scales it is the best understood, as both empirically and (at least in certain limits) theoretically, it is well characterized. Given knowledge of, or an assumption about, the structure of $\overline{\phi}$ within the ABL, V follows directly from surface layer similarity. For most practical purposes this theory can be used to define the surface exchange velocity V:

$$V = C||\hat{\mathbf{u}}||,\tag{14}$$

where C is a dimensionless parameter that depends on the structure and stability of the surface layer [11,14]. For a reasonably wide range of conditions it may be effectively modeled by a constant. Here after we take its value to be 0.0011.

3.1.2 The Mass-Flux, M

From Eq. 10, M defines the turbulent flux just above the ABL. Typical descriptions of the ABL (including, with the exception of our discussion in §5.1, ours) associate h with the depth of the surface-bounded turbulent layer, in which case the presence of a turbulent flux above the ABL appears paradoxical. However if we define the ABL to be the layer which is globally turbulent, then one can admit local, or compact, regions of turbulence associated with the ejection of mass from the ABL. Cumulus clouds, whose active area typically covers a small fraction (2-5%) of the ABL, are good examples of such regions. Typically they are modeled as compact regions (effectively plumes) that loft ABL air into the free atmosphere with some mass (volume) flux M, from which the definition in (10) follows naturally. A more formal derivation of (10) as a model for the flux just above the ABL associated with cumulus clouds is given in [37]. Although this justifies the form of the closure in (10) it says nothing about the value of M. Indeed, how to determine M is a topic of active research and is discussed in more detail in §5.2.

3.1.3 Entrainment, E

Using the definition of \mathcal{D} one can rewrite (10) as

$$(\partial_t + \overline{\mathbf{u}}_+ \cdot \nabla) h_+ = \overline{w}_+ - M + E. \tag{15}$$

This equation states that the change in the height of the interface h, following its motion, is determined by the flux of fluid into the control volume it bounds (given by \overline{w}_+) minus the mass flux out of the volume (in association with cumulus venting and measured by M), and E. It is in this sense that E represents the *diabatic* growth rate of the layer, i.e., the rate at which the turbulent ABL deepens by mixing into the overlying fluid. Hence its identification with entrainment. Another way to look at E is by integrating (8) between h_- and h_+ , in which case for $\phi_- = \hat{\phi}$

$$E = -\frac{\overline{w'\phi'}_{-} + (F_{\phi,+} - F_{\phi,-})}{\phi_{+} - \phi_{-}} \approx -\frac{\overline{w'\phi'}_{-}}{\Delta_{+} - \phi},\tag{16}$$

where the latter equality follows in the case where $\widehat{\phi} = \phi_{-}$ and F is continuous across the interface—in some cases this second approximation is not appropriate [26]. Nonetheless, the approximation serves to place E in a context similar to M and V in that it models an interfacial flux, which in this case is called the entrainment flux. Differences between E and M determine the evolution of the strength of the contact discontinuity at h, i.e., $\partial_t \Delta_+ \phi$.

3.2 Fixed boundary layer depth

In lieu of defining a mass budget for the ABL, and directly modeling the exchanges of mass, heat and moisture between the ABL and its environment, it often proves convenient to simply define a layer of fixed depth to be the model ABL. Typically this may be associated with the lowest model level. In this case integrating (8) over the specified layer results in an equation for its evolution which is identical to (9) ohne terms involving variations in h,

$$h\frac{D\widehat{\phi}}{Dt} - \Delta_{+}\phi \mathcal{D}h = -\Delta \overline{w'\phi'} - \Delta F_{\phi}$$
(17)

In the absence of a mass flux out of the ABL, i.e., $\overline{w'\phi'}_{+} = 0$,

$$h\frac{D\widehat{\phi}}{Dt} = \mathcal{D}h\Delta_{+}\phi + \overline{w'\phi'}_{0} - \Delta F_{\phi},\tag{18}$$

At first glance (18) appears to not involve mixing between the free troposphere and the ABL. However, this mixing is implicit in the first term on the rhs, as for fixed h the mass budget implies $E = \mathcal{D}h$ when M = 0. In this limit, the mixing between the ABL and free troposphere is one way, i.e., the boundary layer is incapable of modifying the free-tropospheric state. The M = 0 limit is problematic for $\mathcal{D} < 0$, as in this case there is no process whereby a steady-state ABL depth can be expected. However, in such cases one would expect a mass flux out of the ABL, choosing this mass flux consistent with the assumption of fixed ABL depth requires:

$$h\frac{D\widehat{\phi}}{Dt} = -\left(\mathcal{D}h + M\right)\Delta_{+}\phi + \overline{w'\phi'}_{0}, -\Delta F_{\phi}.$$
(19)

In the non-entraining limit, with $\mathcal{D} > 0$, $M = -\mathcal{D}h$ and the first term on the rhs of (19) vanishes.

Equation (19) is identical, in form, to what one gets by considering the budget of $\overline{\phi}$ over a layer of fixed depth in the presence of a model of $\overline{w'\phi'}$ in terms of a turbulent diffusivity, K, i.e.,

$$\overline{w'\phi'} = -K\partial_z\overline{\phi}, \quad \text{with} \quad K = \ell M$$
 (20)

where ℓ is a length-scale implicit in the definition of the derivative. To maintain consistency with the implied ABL mass balance the adjustment of the free-troposphere temperature profile through a mass-flux out of the ABL (for instance through a cumulus adjustment scheme) should in the limit of a steady-state ABL depth be balanced by commensurate amounts of entrainment mixing, e.g., per Eq. (19). This emphasizes the need for models employing flux-gradient relations of the form (20), and which wish to rationalize their exchange rules based on our current understanding of boundary layer processes, to consider M in the specification of K. Moreover, given some K it becomes redundant to additionally specify E.

3.3 Energetics

A starting point for a consideration of the bulk energetics is the equation for the turbulence kinetic energy defined here as:

$$e \equiv \frac{1}{2} \left(\overline{u'^2} + \overline{v'^2} + \overline{w'^2} \right). \tag{21}$$

From the governing equations for the velocity field it is straightforward to show that for the case when departures from mean-flow homogeneity do not contribute to the shear production of turbulence (i.e., terms involving $\nabla \overline{\mathbf{u}}$ and $\partial_z \overline{w}, \nabla \overline{w}$)

$$\partial_t e + \overline{\mathbf{u}} \cdot \nabla e + \overline{w} \partial_z e = -\overline{\mathbf{u}'w'} \cdot \partial_z \mathbf{u} + \mathcal{B} - \partial_z \left[\frac{1}{\rho_0} \overline{w'p'} + \overline{w'\mathcal{E}'^2} - \nu \partial_z e \right] - \varepsilon, \tag{22}$$

where $e' = \frac{1}{2} \left(u'^2 + v'^2 + w'^2 \right)$, ε is molecular dissipation and $\mathcal{B} \equiv \overline{w'b'} = g\overline{w'\theta'_v}/\Theta$ defines the buoyancy flux. The first term on the rhs represents the conversion of mean-flow kinetic energy in to turbulence

kinetic energy, \mathcal{B} measures the buoyant production of e at the expense of potential energy, or (if negative) the consumption of e in favor of potential energy. The remaining, three component, term describes various conservative forces which, barring contributions at the boundaries of the flow, simply serve to redistribute turbulence kinetic energy.

Insight into the energetics of simple buoyancy driven layers can be obtained by integrating (22) over the depth of the boundary layer and assuming that the fluxes of e vanish at the surface and h_+ and that the bounding surface do not do work on the ABL, in which case

$$\frac{D}{Dt}\widehat{e} = \widehat{\mathcal{B}} - \widehat{\varepsilon} - \frac{E}{h}\widehat{e}. \tag{23}$$

For the case when h is the only length-scale, one can posit the existence of a velocity scale, w_* , which non-dimensionalizes the dissipation, such that $\varepsilon h/w_*^3$ is universal. In equilibrium (23) implies that in the limit $E/w_* \ll 1$,

$$w_*^3 \propto h\widehat{\mathcal{B}}.$$
 (24)

Such a scaling can also be justified using arguments from the theory of homogeneous isotropic turbulence, in which case w_* simply measures the strength of eddies of the scale h assuming that the turbulence in the boundary layer scales following Kolmogorov's inertial subrange scaling for dissipation rates equal to $\hat{\varepsilon}$. Evaluation of $\hat{\mathcal{B}}$ requires a determination of the profile of \mathcal{B} . From the definition of θ_v , it can be linearly related to fluxes of state variables, i.e.,

$$\mathcal{B} = \left. \partial_s b \right|_{p,q} \overline{w's'} + \left. \partial_q b \right|_{p,s} \overline{w'q'} \tag{25}$$

where the partial derivatives, $\partial_s b$, $\partial_q b$ take on very different values³ depending on whether or not the layer is saturated [9,35]. Thus knowledge of the profile of \mathcal{B} depends on the profiles of $\overline{w's'}$ and $\overline{w'q'}$ and the state of the system (e.g., to determine whether the layer is saturated or not at a particular level). In the absence of sources, and in steady state the flux profiles are just constant. However, for many problems the assumption of stationarity is too strong, and a weaker constraint, called quasi-stationarity is employed. A quasi-steady layer is defined to be one satisfying

$$\partial_t(\partial_z\overline{\phi}) = 0. \tag{26}$$

Differentiating (8) by z this constraint, which says that the shape of the expected profiles is constant in time, is equivalent to requiring $\overline{w'\phi'} + \overline{F}_{\phi}$ to be linear with z. Thus given \overline{F}_{ϕ} and the boundary fluxes, the profile of \mathcal{B} , and hence the energetics of the layer, are straightforward to calculate.

Quasi-steadiness can be expected to hold so long as the boundary forcings are changing on timescales much longer than a turbulent timescale, $t_* = h/w_*$. Because w_* is typically near unity, for an ABL depth of $h \approx 600$ m $t_* \approx 10$ min and hence quasi-steadiness is often a useful constraint. Applying it to a convective ABL for which it is typically assumed that $\mathcal{B}_- = -\kappa \mathcal{B}_0$, quasi-steadiness allows us to evaluate Eq. 24, from which it follows that by setting the constant of proportionality to $2/(1-\kappa)$

$$w_* = (\mathcal{B}_0 h_+)^{1/3} \tag{27}$$

which defines w_* as the velocity scale first introduced by Deardorff [10] to scale the dry convective boundary layer. For this reason (24) with a constant of proportionality of $2/(1-\kappa)$ is often used as a generalized convective velocity scale; which at times is questionable because the underlying assumption of h being the only independent length-scale is not generally justified. For instance in stratocumulus for which \mathcal{B} is discontinuous across $z=\eta$ the cloud base height appears naturally as an additional length-scale [39].

³ Formally b the buoyancy is a function of the state (p, q, s) of the system, in which case the discontinuity of the partial derivatives $\partial_q b, \partial_s b$ is simply the definition of a first order phase transition.

4 Stratocumulus

Many of the assumptions which allowed us to derive (11-13) follow straightforwardly from the assumption that profiles of thermodynamic quantities are well mixed, i.e., $\overline{\phi} = \widehat{\phi}$ for $z \leq h$. The well mixed assumption also justifies the use of surface-layer similarity theory to specify the surface exchange coefficient C in term known quantities, and clarifies the relationship between M and E. An assumption about the vertical profile of the state variables also allows one to diagnose where clouds are to be expected, as well as their properties, which is necessary to compute radiative profiles, and hence \overline{F}_s , as well as the energetics of the layer. Although the assumption of well mixedness does not change the form of (11)-(13), it does make it easier to rationalize the closure assumptions necessary to determine \overline{F}_s , V and E and hence the application of the bulk theory.

The dynamics of stratocumulus topped mixed-layer system are relatively simple. To provide guidance for their incorporation into simplified dynamical models, as well as illustrate their basic behavior we first construct a minimal representation of stratocumulus, one for which the flow is horizontally homogeneous (except at the level of the divergence, i.e., \mathcal{D} is not assumed to vanish), the mass-flux, M, vanishes, and the cloud is non-precipitating, so F_q vanishes. This allow the radiative driving to be specified unambiguously by the symbol \overline{F} . Although ΔF is related to the structure of the ABL, for our immediate purposes we can consider it constant. For sufficiently thick clouds (more than 100 m deep) at night this is actually not a particularly limiting assumption.

With these assumptions the mixed-layer equations can be written as a system of three ODEs:

$$\frac{dh}{dt} = E - \mathcal{D}h,\tag{28}$$

$$\frac{d\widehat{s}}{dt} = \frac{E\Delta_{+}s - V\Delta_{0}s - \Delta F}{h}, \text{ and}$$

$$\frac{d\widehat{q}}{dt} = \frac{E\Delta_{+}q - V\Delta_{0}q}{h}.$$
(29)

$$\frac{d\hat{q}}{dt} = \frac{E\Delta_{+}q - V\Delta_{0}q}{b}.$$
(30)

A specification of the large-scale state can be interpreted to mean a specification of V, \mathcal{D} and the boundary values s_+, q_+, s_0 , and q_0 in which case all that remains to close the system is a specification of E. For now we consider a simple, physically plausible, relation which simplifies the analysis, namely the case when $E = \alpha \Delta F/\Delta_+ s$, where α is a constant of order unity. Heuristically this expression for E suggests that diabatic mixing is proportional to the rate of driving of the flow (as measured by ΔF) and inversely proportional to the stability of the interface at cloud top, as measured by $\Delta_+ s$. Closures which more faithfully respect the energetics of the system are evaluated in [40,44] as well as in §4.4.

4.1 Equilibria

With this model of E we can solve for steady states of the system (which we denote by subscript ' ∞ ')

$$h_{\infty} = h_* \left(\frac{\alpha \sigma}{1 + \sigma - \alpha} \right), \tag{31}$$

$$s_{\infty} = s_0 - \Delta s \left(\frac{1-\alpha}{\sigma}\right), \text{ and}$$
 (32)

$$q_{\infty} = q_0 + \Delta q \left(\frac{\alpha}{1+\sigma}\right),\tag{33}$$

where

$$h_* = \frac{\Delta F}{D\Delta s}$$
 and $\sigma = \frac{V\Delta s}{\Delta F}$. (34)

The height-scale h_* measures the height the equilibrium boundary layer would obtain for $\alpha = 1$, wherein h becomes independent of σ . In the stratocumulus regions of the northeast Pacific typical parameter values are $\mathcal{D} \approx 4 \times 10^{-6} \text{ s}^{-1}$, $V \approx 0.008 \text{ m s}^{-1}$, $\Delta s \approx 18 \text{ kJ kg}^{-1}$, and $\Delta F \approx 40 \text{ W m}^{-2}$ averaged over a diurnal cycle, yielding values of $h_{\infty} \approx 800 \text{ m}$ and $\sigma \approx 2.5$ [42]. In this respect it is worth noting that cloudy solutions driven with diurnally varying radiative forcing and averaged over the diurnal cycle do not differ appreciably from equilibrium solutions driven by diurnally averaged radiative forcings [48].

An interesting aspect of the equilibrium equations is that for reasonable parameter values they predict realistic equilibrium ABL depths. They also show that for increasing lower tropospheric stability, as measured by increasing σ , $q_{\infty} \to q_0$. Because $s_{\infty} < s_0$ for $\alpha < 1$, one expects moister ABLs, and hence lower cloud bases with increasing lower-tropospheric stability, which is consistent with the climatological data [15].

One can help fix these ideas by looking at the height of the cloud base η . We can solve for η_{∞} by recognizing that η is just the height z where $q_s(T(s,z))$ (the saturation vapor specific humidity) is equal to q. Similarly, η_{∞} is just the height z where $q_s(T(s_{\infty},z)) = q_{\infty}$. To solve for z we need a tractable form for q_s , and an expression for T in terms of s and z. The latter follows directly from the definition of s recognizing that $q_l = 0$ at cloud base; the former we obtain by linearizing the logarithm of the dependence of saturation vapor pressure about the surface temperature such that

$$q_s(T) = q_s(T_0) \exp\left[\frac{L_v}{R_v T_0^2} (T - T_0)\right]. \tag{35}$$

Noting that because the lower surface is saturated, i.e., $q_s(T_0) = q_0$,

$$q_{\infty} = q_0 \exp\left(-\frac{\Delta F}{g\eta_* V}(1-\alpha)\right) \exp\left(\frac{-\eta_{\infty}}{\eta_*}\right) \quad \text{where} \quad \eta_* = \frac{R_v s_0^2}{L_v c_p g} \approx 1500 \text{m}. \tag{36}$$

it is straightforward to arrive at the following explicit equation for η_{∞} :

$$\eta_{\infty} = -\eta_* \ln \left(1 + \frac{\alpha}{1+\sigma} \frac{\Delta q}{q_0} \right) - \frac{\Delta F}{Vq} (1-\alpha). \tag{37}$$

Physically, the last term on the right arises because the temperature just above the surface is colder then the surface when $\alpha < 1$. This leads to lifting condensation levels being a few hundred meters lower than one would expect were they to assume that $s_{\infty} = s_0$. In the sub-tropics, q_+ is very small compared to q_0 , so that $\Delta q/q_0 \approx -1$. With this approximation, in the large σ limit,

$$\eta_{\infty} \to \eta_* \frac{\alpha}{1+\sigma} - \frac{\Delta F}{Vq} (1-\alpha),$$
(38)

from which it is readily apparent that both larger σ and smaller α favor the lowering of cloud base.

Of course one should hesitate to associate small η with a high-probability of cloud incidence. A better measure would be the ratio η/h . From (31) and (37) in equilibrium this ratio is,

$$\frac{\eta_{\infty}}{h_{\infty}} = -\frac{\eta_{*}}{h_{*}} \left[\frac{\ln\left(1 + \frac{\alpha}{1+\sigma} \frac{\Delta q}{q_{0}}\right)}{\sigma \alpha} \left(1 + \sigma - \alpha\right) \right] - \frac{\Delta F}{Vgh_{*}} \left[\frac{(1-\alpha)(1+\sigma-\alpha)}{\sigma \alpha} \right]$$
(39)

Equation (39) is plotted in Fig. 2. Cloudy equilibria are expected when $\eta_{\infty}/h_{\infty} < 1$. The figure shows that this ratio decreases markedly with decreasing α , and is relatively insensitive to σ . This tells us that as entrainment becomes less efficient, the non-dimensional thickness of the cloud layer increases. This makes physical sense. The free troposphere tends to be both warmer and drier than the ABL, thus enhanced entrainment can be expected to increase the lifting condensation level more rapidly than in increases h. The marked sensitivity of the solutions to α motivates a tremendous amount of research in this area. [20–22,17,26,25,40]

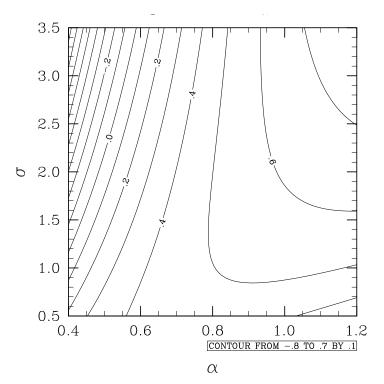


Fig. 2 Equilibrium cloud base height normalized by ABL depth as a function of α and σ .

4.2 Adjustment

The governing systems of equations are weakly non-linear due to the inverse dependence on h in (29) and (30) and from the dependence of $\Delta_+ s$ on \hat{s} in the model of E. Linearizing (28)-(30) about its equilibrium state effectively amounts to fixing $E = E_{\infty}$ in (28) and (30) and setting $h = h_{\infty}$ in the denominator of Eqs. (29) and (30). Analysis of this linearized system yields three real eigenvalues,

$$\lambda_1 = -\mathcal{D}, \quad \lambda_2 = -\mathcal{D}\left(\frac{1+\sigma-\alpha}{\alpha}\right), \quad \text{and} \quad \lambda_3 = -\mathcal{D}\left(\frac{(\sigma+\alpha)(1+\sigma-\alpha)}{\alpha\sigma}\right)$$
 (40)

which are all negative if $\mathcal{D} > 0$ and $\alpha < 1 + \sigma$, thus defining the domain of stable solutions. In the more general case, when $\alpha = \alpha(h)$, stability demands that $\partial_h \alpha < \mathcal{D}$. Such an analysis could be similarly extended to the more realistic case, when $\alpha = \alpha(s, q, h)$.

Roughly speaking the inverse of these eigenvalues corresponds to the timescale on which perturbations in h, s and q are damped, which for typical parameter values evaluates to roughly 70, 30 and 20 hrs respectively. On such timescales, one expects the advective terms in the substantial derivatives, e.g., $\hat{\mathbf{u}} \cdot \nabla \phi$, to be significant [35,36], which motivates their inclusion in our subsequent analysis below.

4.3 Multiple-Equilibria

The systems we have considered have only a single equilibrium. However, when one allows for the radiative forcing, ΔF to depend on the cloudiness, and for other processes, such as turbulence generated by shear at the surface, to generate entrainment it is possible for the system to have multiple equilibria and exhibit hysteresis effects. An example of this behavior was shown by [31] using a simple model of shear generated turbulence. The multiple equilibria for their model consisted of a shallow cloud free ABL, and a deeper cloudy ABL.

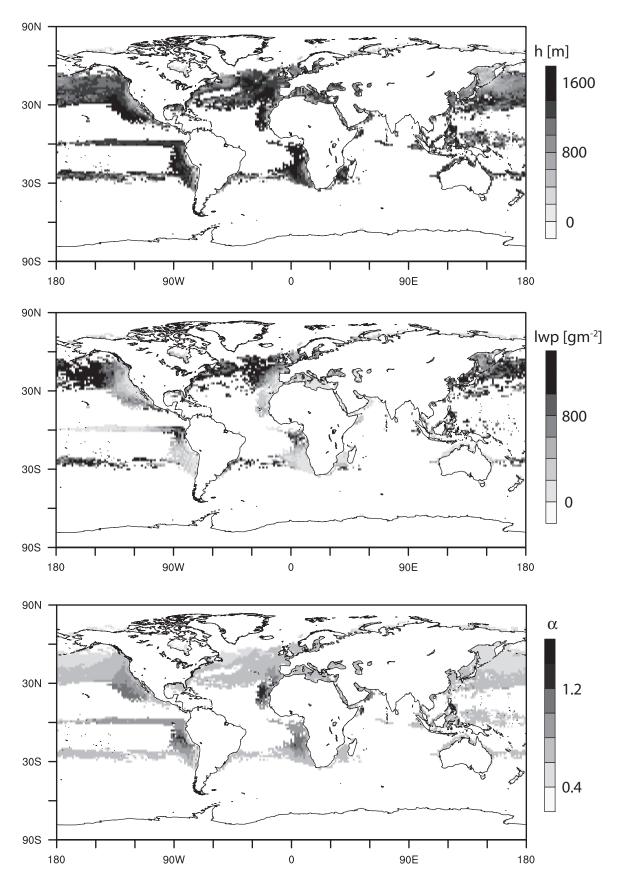


Fig. 3 Equilibrium solutions using ERA40 climatological forcing for June, July and August, with $\alpha_{\mathcal{B}}=0.25$, constant advective forcings and fixed radiative cooling representative of mean stratocumulus regions. From top to bottom, h, liquid water path \mathcal{L} , and α . Values with $\mathcal{L} > 1400$ g m⁻², or h > 2000 as masked.

4.4 Realistic entrainment formulations

To qualitatively evaluate the utility of equilibrium solutions of the mixed layer theory, in Fig. 3 we show solutions computed using the June, July and August 1983-2001 climatology as extracted from the 40 plus year reanalysis of meteorological data by the European Center for Medium Range Weather Forecasts [46, e.g., the ERA40]. The climatology was used to specify the surface and free tropospheric state, with the divergence of the near-surface winds being substituted for \mathcal{D} . In computing these solutions, the radiative forcing was fixed at 40 Wm⁻² and advective tendencies ($\nabla q = 1.2 \times 10^{-9} \text{ m}^{-1}$, $\nabla s = 3.9 \times 10^{-3} \text{ m s}^{-2}$ and $\nabla h = 1.4 \times 10^{-4}$) were added based on values characteristic of the mean stratocumulus regions of the northeast Pacific, as derived from the ERA40, using their representation of ABL depth.

These solutions were also computed using a state of the art entrainment parameterization, for which the entrainment rate is determined by requiring the net buoyancy flux to be some fixed fraction of its value in the absence of entrainment [17]. Mathematically the closure amounts to finding which value of E results in a predetermined value of the ratio:

$$\alpha_{\mathcal{B}} \equiv \widehat{\mathcal{B}}/\widehat{\mathcal{B}}_{E=0},\tag{41}$$

where $\widehat{\mathcal{B}}_{E=0}$ denotes the average buoyancy flux in the absence of entrainment. Although straightforward, the evaluation of $\widehat{\mathcal{B}}$ and $\widehat{\mathcal{B}}_{E=0}$ is not trivial. E is determined by this ratio because E helps define the thermodynamic fluxes at cloud top, which are then related to \mathcal{B} given the assumption of quasisteadiness and the piecewise continuous relations in Eq. (25). This fixed $\alpha_{\mathcal{B}}$ closure is similar in spirit to the fixed α closure discussed earlier, but by integrating the actual and zero entrainment buoyancy fluxes, it more accurately measures the energy available to drive entrainment, and the energetic cost of entrainment. It is chosen because it is conceptually simple, yet has faired well (with $\alpha_{\mathcal{B}} \approx 0.2$) in past evaluations based on both large-eddy simulation and field data [44].

The solutions in Fig. 3 plausibly represent the well known regions of stratocumulus convection, over the eastern boundary currents of the subtropical oceans. In qualitative agreement with climatology four main regions appear [15]. In the NE Pacific near 120 W and 30 N; in the SE Pacific near 80 W and 20 S; in the N. Atlantic near 20 W and 30 N, and in the S. Atlantic near 10 E and 20 S. Over each of these regions the equilibria are characterized by shallow ABLs between 400 and 1000 m topped by relatively thin stratocumulus layers, with liquid water paths ($\mathcal{L} \equiv \int_0^h \rho_0 q_l dz$) generally less than 400 g m⁻². Because deviations in the well mixed assumption which are small from the perspective of the budgets of q and s can have a large impact on the equilibrium value of q_l , we relax the well mixed assumption when calculating the liquid water path to two thirds of its adiabatic value, making it more commensurate with values commonly observed [42].

The model also produces equilibria over the storm tracks in the northern hemisphere and along thin bands over the equatorial cold tongue of the Pacific and over both the Pacific and Atlantic oceans near 25 S. Although these also tend to be regions of significant low cloudiness the equilibrium solutions in these areas tend to have such thick clouds that precipitation and other processes will become important—suggesting that the stratocumulus regions may be well captured by the model after being masked for equilibria whose cloud depths are incompatible with the assumption of no, or little, precipitation.

Another basis for flagging solutions of the model is by measuring the value of $\alpha = E\Delta_+ s/\Delta F$ the fixed $\alpha_{\mathcal{B}}$ parameterization produces in equilibria. Values of $\alpha > 1$ are indicative of other processes (surfaced and latent heat fluxes) contributing significantly to the energetics of the layer. In a remarkably insightful study Bretherton and Wyant [9] show that in this limit the amount of entrainment warming can be larger than the radiative cooling of cloud top air parcels, and thus while evaporation of liquid water may fuel their descent through the cloud layer, they tend to be inhibited from mixing to the surface, thus resulting in the thermodynamic differentiation between the cloud and sub-cloud layer and more trade-cumulus like cloud circulations. Formally this transition can be associated with buoyancy fluxes just below cloud base reaching the limit associated with the dry convective ABL, i.e., $\mathcal{B}_{\eta} = -\kappa \mathcal{B}_{0}$, [39] which in practice often amounts to $\alpha > 1.1$ or so [45]. So although Fig. 3 is masked based on values of h > 2000, these regions are invariably associated with $\alpha > 1.1$ and thus could be just as effectively masked by this energetic constraint.

In either case, the transition of the equilibria to higher values of α as the flow advects westward and equatorward away from the eastern-boundary currents of the subtropical oceans is indicative of

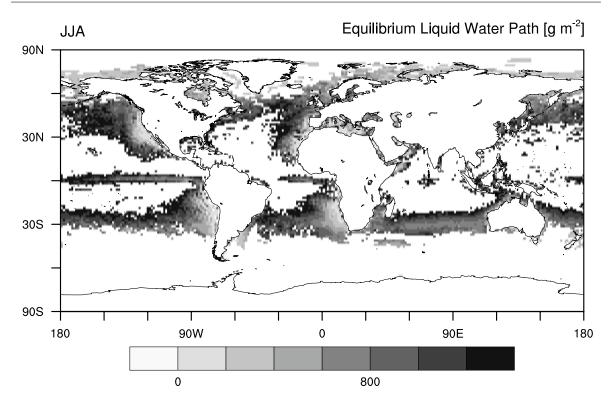


Fig. 4 Equilibrium solutions for $\alpha = 0.85$ and boundary and forcings as in Fig. 3.

a greater propensity toward lower cloud fractions and more trade-wind like circulations, consistent with observations of this evolution. More quantitative evaluations of this approach will be presented elsewhere, and although not shown, the seasonal cycle is also qualitatively captured by this approach, with cloudiness in the southern oceans peaking in September-October-November as observed. Moreover, the inclusion of more realistic (spatially varying) advective effects leads to further improvement in the solutions.

For the sake of applying these types of approaches to modeling the stratocumulus topped ABL in very simple models of larger-scale circulations we note α does not vary tremendously over the solution region in Fig. 3. Which suggests that the much simpler, fixed α , model may have utility. These are evaluated in terms of the liquid water path for $\alpha = 0.85$ in Fig. 4, here again the area of modest liquid water paths ($< 500 \text{ g m}^{-2}$) well inscribe the stratocumulus regions. Because such a model permits algebraic solutions, but is formally related to much more sophisticated representations of stratocumulus, its equilibrium solutions may provide a useful alternative to purely empirical parameterizations of stratocumulus currently used in many models of simplified tropical dynamics.

5 Trade-wind cumulus

In contrast to stratocumulus, the trade-wind layer is less obviously a single layer: past work has emphasized distinctions between the cloud and sub-cloud layer, as well as the mediating influence of thin layers near cloud base (the transition layer) at the surface, and at the top of the cloud layer (e.g., Fig. 1 in [27]). As a result a variety of bulk models have been developed to represent such effects. These various approaches are discussed in turn below.

5.1 One-layer models

A well known bulk model of the trade-wind regime is that developed by Betts and Ridgway [7]. The model incorporates the entirety of the trade-wind layer, from the surface to the top of the cloud layer,

into a single layer Choosing h to represent the entirety of the layer avoids the problem of dealing with a model of the mass flux and so, as with stratocumulus, the system can be represented by (11)-(13) with M=0.

By considering only non-advective steady-state solutions Betts and Ridgway substitute for E in (12) and (13) with $\mathcal{D}h$ as implied by the steady state mass balance. In this case the heat and moisture budgets become

$$0 = \mathcal{D}h\Delta_{+}s + \overline{w's'}_{0} - \Delta F \tag{42}$$

$$0 = \mathcal{D}h\Delta_{+}q + \overline{w'q'}_{0}. \tag{43}$$

The model amounts to an exploration of the above constraints under the assumption that: (i) the vertical structure of the state variables within the ABL is well described by a prescribed form (a mixing line [6]); and (ii) that the surface flux $\overline{w'\phi'}_0$ (for $\phi \in \{s,q\}$) is well modeled by (10) with V determined by surface layer similarity and $\widehat{\phi}$ replaced by ϕ_m in $\Delta_0 \phi$, where subscript m denotes a value in the sub-cloud mixed layer.

To illustrate the behavior of their model we specify an analogous vertical structure as follows:

$$\overline{\phi} = \begin{cases} \phi_m & z < \eta \\ \phi_m + \left(\frac{z-\eta}{h-\eta}\right)\phi_- & z \ge \eta \end{cases} \quad \text{where} \quad \phi_- = (1-\beta_-)\phi_m + \beta_-\phi_+; \tag{44}$$

and recall that η denotes the height of the lifting condensation level and thus measures the depth of the sub-cloud layer, while $\phi_- \equiv \overline{\phi}_{z=h_-}$ measures the state variable properties just below cloud top and is consistent with the cloud-layer lying on a mixing line defined by the parameter $\beta_- < 1$. Integrating (44) over the layer yields

$$\widehat{\phi} = \phi_m + \zeta(\phi_+ - \phi_m) \quad \text{where} \quad \zeta = \frac{\beta_-}{2} \left(1 - \frac{\eta}{h} \right) \tag{45}$$

where ζ is a non-dimensional height⁴. This definition of the vertical structure allows us to write the heat and moisture budgets, i.e., (42) and (43), as

$$0 = \mathcal{D}h(1-\zeta)(s_{+}-s_{m}) + V(s_{0}-s_{m}) - \Delta F, \tag{46}$$

$$0 = \mathcal{D}h(1-\zeta)(q_{+}-q_{m}) + V(q_{0}-q_{m}). \tag{47}$$

Because s_+ is known only as a function of h, after specifying $\mathcal{D}, V, q_+, \zeta$ and the SST, (46) and (47) retain three unknowns. Thus an additional constraint is required to close the system. This is provided by requiring the heat budget (42) to also be balanced over the sub-cloud layer with

$$\mathcal{D}\eta(\overline{s}_{z=n+\epsilon} - s_m) = -\kappa V(s_0 - s_m),\tag{48}$$

i.e., the heat flux at the top of the sub-cloud layer is some fraction $-\kappa$ of its surface value. Specifying $\kappa=0.25$ [7, in analogy to the well established behavior of the dry convective ABL] determines s_m given as a function of the radiative flux divergence across the sub-cloud layer, which we denote by $\Delta_m F$. As was the case for stratocumulus the model is closed by an entrainment assumption, but in this case the assumption is enforced at cloud base, rather than at h.

Substituting s_m into (46) and rearranging (47) allows us to write the heat and moisture budget constraints as

$$s_m = s_0 \left(1 - \frac{\gamma}{1 + \kappa} \right),\tag{49}$$

$$q_m = \frac{q_0 + Rq_+}{1 + R}$$
, and (50)

$$s_{+} = s_{0} \left[1 - \frac{\gamma}{R} \left(\frac{1+R}{1+\kappa} + N \right) \right] \tag{51}$$

⁴ Betts and Ridgway denote this non-dimensional height by α , we use ζ because α is reserved for the stratocumulus entrainment efficiency.

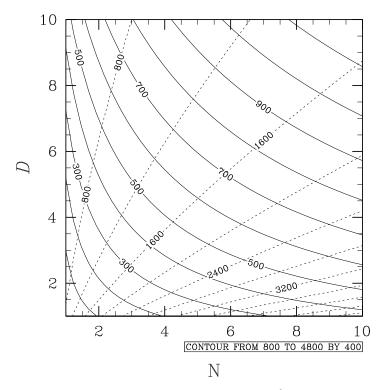


Fig. 5 Equilibrium η and h (dashed) as a function of N and $\mathcal{D} \times 10^6$ for the parameter values in Table 1

where ideally the non-dimensional parameters

$$\gamma = \frac{\Delta_m F}{s_0 V} \quad N = \frac{\Delta F}{\Delta_m F} \quad R \equiv \frac{\mathcal{D}h}{V} (1 - \zeta) \tag{52}$$

help highlight, rather than obscure, the structure of the model. With values typical for the trades, i.e., $\rho\Delta_mF=12~\mathrm{Wm^{-2}},~\Delta F=50~\mathrm{W~m^{-2}},~V=0.01~\mathrm{m~s^{-1}},~\mathrm{and}~\mathcal{D}=3.5\times10^{-6}~\mathrm{s^{-1}},$ the above parameters are roughly 1/300, 5, and 0.5 respectively.

Because both R and the radiative fluxes (and hence N and γ) depend on the structure of the ABL, Eqs. (49)-(51) implicitly define the state of the system and must be solved iteratively. Eq. (51) is actually an equation for the ABL depth, as h is determined by finding the height at which the specified free-atmospheric profile \bar{s}_f matches s_+ . Following the derivation leading to (37), given s_m and q_m , the lifting condensation level follows naturally as:

$$\eta_{\infty} = -\eta_* \ln \left[1 + \frac{\Delta q}{q_0} \left(\frac{R}{1+R} \right) \right] - \frac{\gamma}{1+\kappa} \left(\frac{s_0}{g} \right). \tag{53}$$

Betts and Ridgway solve the above system by using an elaboration on the assumed ABL structure (to account for differences between cloudy and non-cloud regions presumed to exist in fixed proportions) in conjunction with a radiative transfer model to determine F and by specifying the product $\mathcal{D}h(1-\zeta)$ which appears in the definition of R. Doing so obscures the dependence of the solutions on \mathcal{D} and can yield solutions which are inconsistent with the assumed structure of the model, which in our case is determined by β_- and (44). We illustrate the behavior of the model in Fig. 5 by instead specifying the radiative fluxes and \mathcal{D} , with a consistent value of R determined by iteration. So doing can generate an inconsistency between the assumed structure and the specified changes in the radiative fluxes across the layer, but more clearly illustrates the behavior the the model for an externally specified forcing. In calculating these solutions we specify the free atmospheric profile of s so that $s_+/c_p=291.4+0.0058h$, which is based on the mean structure of the atmosphere in the summertime trades of the northeast Pacific. Other parameter values are typical and chosen following Betts [7] and are listed in Table 1. From Fig. 5 with N=5 and $\mathcal{D}=3.5\times10^{-6}~\mathrm{s}^{-1}~\eta\approx600~\mathrm{m}$ and $h\approx1750$. Values of η between 400 and

Table 1	Parameter	values	for use	in	the	construction	of Fig.	5
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Parameter	Value
$\Delta_m F$ SST	$12~{ m W}~{ m m}^{-2}$ $300~{ m K}$
q_+	3 g kg^{-1}
$V \\ p_0$	0.01 m s^{-1} 1013 hPa
κ	0.25
eta	0.3

800 m and h between 1200 and 3000 m are realistic. The tendency for h to be more variable than η reflects the fact that the latter is more sensitive to the dynamical state of the atmosphere as represented by \mathcal{D} . As an aside, our calculations also confirm that the solutions are relatively independent of β_{-} and κ . Choosing $\kappa=0$ results in little change in h but lowers cloud base by roughly 20%; doubling β_{-} also has little effect on h, but lowers cloud base by about 10%. Because for a well mixed sub-cloud layer the height of cloud base, effectively measures the partitioning of the surface fluxes among the sensible and latent components (i.e., the Bowen Ratio) the sensitivity of η to β_{-} and κ measures the extent to which the partitioning of the surface fluxes is controlled by these parameters.

The reasonable value of solutions, their robustness, and their ability to sensibly respond to macroscopic changes in the large-scale state are all attractive elements of this model, and motivates its use in a variety of thermodynamic models of tropical circulations [7,24, among others]. The disadvantage of using this system of equations as a basis for the ABL treatment in simplified models of tropical dynamics, is that h varies considerably—reaching through the depth of the troposphere as the convecting zones are approached (\mathcal{D} decreases). Incorporating the cloud layer into the ABL has advantages from the perspective of continuity with the upstream stratocumulus layers, but disadvantages from the perspective of matching to regions of deeper, precipitating convection. Moreover, as h increases, some of the assumptions of the underlying equations (for instance, neglect of baroclinic effects) also become increasingly suspect. These reasons motivate an exploration of bulk models of just the sub-cloud layer, a topic which we address in §5.2.

A quite different weakness of the (49)-(51) is that the structure of the cloud layer is imposed by an a priori constraint on the structure of the cloud layer (in our case Eq. 44) and the only remaining physics in the model is through the closure assumption (48). Allowing the cloud layer structure to respond to the macroscopic structure of the ABL (i.e., η and h) provides an opportunity for physical insights, and important new feedbacks between the radiative driving of the layer and its bulk structure. Such effects could conceivably be included by letting β_- , or more generally (44), respond more realistically to the bulk state. Another approach to this weakness has been the exploration of two-layer models, which explicitly model the interaction between the cloud and sub-cloud layer, and the cloud layer and the free-troposphere. A well known model in this class is that devised by Albrecht [2] but because of unsatisfactory parameter sensitivities [8] and inconsistencies in its closure [4] this model has not met with general acceptance as a bulk theory of the trade-wind layer.

Before proceeding to other models of the trade-wind layers, we briefly point out some connections between the Betts and Ridgway model of the trade-wind layer and the bulk model of stratocumulus presented in §4. To do so it proves useful to note from (29) and (42) that its equilibrium solutions correspond to an entrainment rate,

$$E = \mathcal{D}h = \frac{1}{(1 - \zeta)(s_{+} - \widehat{s})} \left(\Delta F - \frac{4}{5} \Delta_{m} F \right)$$
(54)

which implies an entrainment efficiency (e.g., §4) of

$$\alpha \equiv \frac{E\Delta_{+}s}{\Delta F} = \frac{1}{(1-\zeta)} \left(1 - \frac{4}{5N} \right). \tag{55}$$

For typical trade-wind situations $N \approx 5$ and $\zeta \approx 0.25$, hence $\alpha \approx 1.1$. In principle this provides a basis for matching equilibrium stratocumulus solutions such as described in §4 to trade-wind solutions.

The definition of the entrainment efficiency (55) allows one to define a dynamic system of equations for the Betts-Ridgway model equivalent to (28)-(30) but replace E with $E + \zeta V$ in (29) and (30) so as to account for the effect of the vertical structure on the surface fluxes. Note that for a well mixed layer $\zeta = 0$ and hence the model reverts to the stratocumulus limit.

5.2 Sub-cloud layer models

As an alternative to a model of a single-bulk layer for the entirety of the trade-wind layer, it can prove convenient to simply consider the ABL as being confined to the sub-cloud layer. As we discussed above, maintaining h near η leads to shallower and less variable values of h over the tropical oceans, more constrained vertical structure, less possibility for baroclinic influences, and greater continuity with periods, or regions, of deeper convection. From such a perspective the governing system of equations are simply (11)-(13). In the stratocumulus limit (e.g., §4) M=0 and ϕ_+ is given the free tropospheric values thereby requiring only one additional constraint to close the system. In the trade-wind regime $M \neq 0$ and ϕ_+ corresponds to values just above cloud base, in the cloud layer, and thus is not-known a priori. Hence closure of this system of equations over the sub-cloud layer in the trades requires four additional constraints. Typically these are applied as models for s_+, q_+ (in effect assuming the structure of the cloud layer) M and E.

An early pioneer of this approach was again Betts [5], who closed the system by writing additional budget equations for the transition layer so as to specify (s_+, q_+) , specifying E following the heat flux closure in (48), and by defining the mass-flux, M implicitly by setting $h = \eta$. Neggers et al., revisit this approach with the aim of developing simple parameterizations of shallow convective ABLs for use in simplified climate models [27]. In their approach they dispense with the transition layer equations and simply fix

$$\phi_{+} = \beta_{\phi_{+}}(\phi_{f} - \widehat{\phi}) \tag{56}$$

for $\phi \in \{s, q\}$, where subscript f refers to the free-tropospheric value which is presumed known. The entrainment flux is once again specified following the well known dry-convective limit, such that

$$E = -\kappa B_0 / \Delta_+ b,\tag{57}$$

with

$$\Delta_{+}b \equiv b_{+} - \hat{b} = \frac{g}{s_{0}} \left[\Delta_{+}s + 0.608s_{0}\Delta_{+}q \right] \quad \text{and} \quad \mathcal{B}_{0} = -\frac{gV}{s_{0}} \left[\Delta_{0}s + 0.608s_{0}\Delta_{0}q \right]$$
(58)

denoting the buoyancy jump at h and the surface buoyancy flux respectively. Lastly the cloud-base mass flux is modeled as

$$M = \mathcal{A}w_* \tag{59}$$

where w_* is the Deardorff velocity scale give by (27) and $\mathcal{A} > 0$ is the cloud fraction determined as a function of the state of the system. Except for (59) these closures are relatively standard: (56) is effectively an assumption on the structure of the cloud layer. Eq. (57), which can be equivalently phrased as having assuming that the entrainment buoyancy flux \mathcal{B}_- is some fixed negative fraction κ of the surface buoyancy flux, is analogous to (48); its application to the buoyancy budget (rather than the heat budget) is standard for the dry convective layer and, notwithstanding arguments to the contrary [7], more appropriate to cumulus topped ABLs as well [39,41,38]. As Neggers *et al.*, point out the form of (59) dates back to much earlier studies, but relating \mathcal{A} to the cloud fraction (especially following their prescription) is novel [27,28].

An interesting aspect of the above model is that it predicts regions of cloud free solutions, effectively those regions where the steady-state h is less than η for M=0. If advective and radiative processes

are negligible compared to mixing processes in the cloud layer then s_+ and q_+ can be expected to lie on a mixing line, in which case $\beta_{s_+} = \beta_{q_+} = \beta_+$, and equilibria take the form

$$h_{\infty} = \frac{E}{\mathcal{D}},\tag{60}$$

$$\widehat{s}_{\infty} = \frac{V s_0 + \beta_+ E s_f - \Delta F}{V + \beta_+ E}, \quad \text{and}$$
(61)

$$\widehat{q}_{\infty} = \frac{Vq_0 + \beta_+ Eq_f - h\widehat{\mathbf{u}} \cdot \nabla q}{V + \beta_+ E},\tag{62}$$

where

$$E = \frac{\kappa}{\beta_{+}} V \left(\frac{\Delta F + 0.608s_{0}h\hat{\mathbf{u}} \cdot \nabla q}{\Delta F_{s} + 0.608s_{0}h\hat{\mathbf{u}} \cdot \nabla q + (1+\kappa)V(\Delta s + 0.608s_{0}\Delta q)} \right).$$
 (63)

Following Neggers et al., the radiative forcing ΔF and the advective tendency in the moisture equation $(h\hat{\mathbf{u}} \cdot \nabla q)$ are retained so as to represent both advective and radiative effects. Given \hat{s} and \hat{q} , it is straightforward to show that

$$\eta_{\infty} = \Delta_0 s/g - \eta_* \ln(\widehat{q}/q_0). \tag{64}$$

Cloud free solutions thus correspond to solutions where $\eta > h$. For $\eta < h$ we expect cloudy solutions. For these we note from (63) that for a fixed forcing⁵, and s_f constant with height, E is independent of h, which implies \widehat{s} and \widehat{q} and hence η are also independent of h. Consequently M can be solved for directly, simply by noting that in this situation h = E/D and that in so far as the mass flux principally acts to ventilate the ABL, thereby relaxing h to η , choosing $M = E - D\eta$ must be true in equilibrium, in which case the equilibrium cloud fraction is just

$$\mathcal{A} = \frac{E - \mathcal{D}\eta_{\infty}}{w_{\star}}.\tag{65}$$

For typical values of $E \approx 1$ cm s⁻¹, $D = 3^{-6}$ s⁻¹, $\eta = 500$ m and $w_* = 0.6$ ms⁻¹, $E \gg \mathcal{D}\eta$, and $\mathcal{A} \approx E/w_* \approx 0.02$ which might explain why cloud fraction is order a few percent.

This special (albeit somewhat contrived) limit of constant forcings and s_f independent with height highlights the role of β_+ . Because $E \propto \beta_+^{-1}$ but the heat and moisture budgets depend only on the product $\beta_+ E$, both \hat{s} and \hat{q} , and hence η must be independent of β_+ . Thus, in this limit, β_+ plays no role in determining the strength of the surface fluxes, or the height of the sub-cloud layer, but instead simply determines the depth of the equilibrium layer whose mass flux is zero—and by implication the mass flux for models of M that act to ventilate the ABL thus relaxing h to η .

Perhaps a more realistic situation is obtained by letting s_f vary with height. In principal, s_f , ΔF , and $h\hat{\mathbf{u}} \cdot \nabla q$ should all depend on the structure of the trade-wind layer, but here we only consider s_f varying with height as $s_f/c_p = 301 + \Gamma h$. Because s_f really is determined by the height of cloud top and h is intended to model the cloud base height, this solution is at best a modest improvement. These limitations not with standing, Fig. 6 shows the M=0 solutions of (60)-(62), for $\beta_{+}=0.2$ and the aforementioned model of s_f , as a function of \mathcal{D} and Γ . Because of the h dependence of s_f they are obtained iteratively. The solid lines show contours of η while the dashed lines show $h-\eta$. Thus cloud free (M=0) solutions are associated with negative dashed contours and the zero contour marks the values of \mathcal{D} and η at which $h = \eta$. If we denote these values of \mathcal{D} by \mathcal{D}_* then for some $\mathcal{D} < \mathcal{D}_*$ the equilibrium mass flux can be inferred as $M = (\mathcal{D}_* - \mathcal{D})\eta$. From Fig. 6 we note that $\mathcal{D}_* \approx 3.5 \times 10^{-5}$ and η varies between values just below 400 m to just above 800 m. Because typical values of \mathcal{D} in the trades are an order of magnitude less than \mathcal{D}_* these results suggest that η increases with decreasing stability and that $M \approx 2.5$ cm s⁻¹, consistent with the results of Neggers et al., [27]. These simple calculations indicate that the model does provide a useful description of the equilibrium state of the sub-cloud layer, and thus could prove beneficial in a variety of contexts as discussed in the introduction. A more thorough evaluation, which complements the analysis of the stratocumulus layer in $\S 4.4$ is provided by

⁵ i.e., s_f , ΔF_s and $h\hat{\mathbf{u}} \cdot \nabla q$ constant.

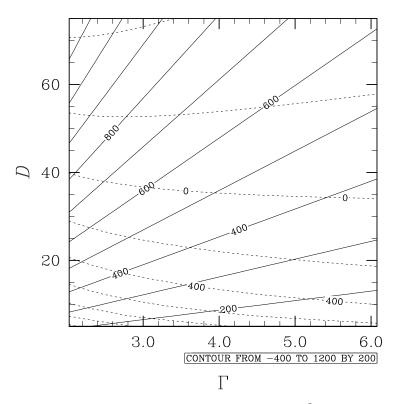


Fig. 6 Equilibrium η (solid) and $h - \eta$ dashed as a function of $\mathcal{D} \times 10^{-6}$ and Γ . with other parameters as specified in Table 1, and ΔF and $h\hat{\mathbf{u}} \cdot \nabla q$ taken as 15 and 30 Wm⁻² respectively, consistent with the values chose by [27] for a 750 m deep sub-cloud layer.

One unsatisfactory aspect of this class of models is their dependence on a seemingly poorly constrained parameter, β_+ . For the solutions in Fig. 6, halving β_+ effectively doubles \mathcal{D}_* and hence the equilibrium mass flux. One possible resolution of this sensitivity is to note that the similarity structure of a dry convective ABL growing into a uniformly stratified fluid which results in the closure assumption $\mathcal{B}_- = -\kappa \mathcal{B}$ also predicts

$$\Delta_{+}b = N_c^2 h \left(\frac{\kappa}{2\kappa + 1}\right) \tag{66}$$

where N_c^2 is the Brunt-Väisälä frequency in the cloud layer. For a moist adiabatic lapse rate of $d\theta_v/dz$ of 5 K km⁻¹ this corresponds to $\Delta_+\theta_v=(\Theta/g)\Delta_+b\approx 0.36$ K for a sub-cloud layer of 500 m and a moist adiabatic cloud layer $(N_c^2=1.6^{-4}~{\rm s}^{-2})$. Given Δ_+b one can then solve for β_+ . Note that from (65), in the limit of $E>>\mathcal{D}$, and with w_* given following (27) this constraint implies that

$$\mathcal{A} = \frac{2\kappa + 1}{Ri} \quad \text{where} \quad Ri \equiv \frac{h^2 N_c^2}{w_*^2},\tag{67}$$

is a bulk Richardson number. For typical parameter values this yields values of \mathcal{A} of about 2%, which is reasonable based on past studies. Preliminary tests of this idea are promising, but it remains to be evaluated more carefully.

6 Discussion

For the purposes of diagnostic studies, or simple thermodynamic models of the tropical atmosphere, coupling simple equilibria of the stratocumulus limit with the trade-wind equilibria of the Betts and Ridgway model ($\S 5$) seems like an attractive approach. In this context h consistently represents the

cloud top, or inversion, and hence the two approaches have a compatible mass budget. Moreover, because the decoupling transition of stratocumulus is formally identified with the cloud base closure assumption (albeit on the buoyancy rather than the heat budget) used by one-layer bulk models the possibility of matching solutions by choosing the β_{-} representation of the trade-wind layer which best matches the mass budgets of the two models at the transition point seems natural. One avenue of future exploration would be to explore such matching in the context of the existing climatological boundary forcings (such as provided by modern reanalyses) and realistic representations of radiative forcings. In so doing it would seem worthwhile to incorporate more sophisticated representations of partial cloudiness. In this respect the random mixing line model developed by Park [29] appears to be a promising approach.

For the purposes of defining a shallow ABL for inclusion in dynamic models it probably makes more sense to couple the sub-cloud model of §5.2 with the stratocumulus model of §4. Here, the mass budgets of the two layers are not strictly compatible, with h formally including a thin cloud layer for the stratocumulus limit, and excluding the cloud layer in the sub-cloud layer limit. So while the matching condition at the transition point does maintain continuity in the sub-cloud buoyancy flux, i.e., sub-cloud solutions are sought when stratocumulus solutions predict $\mathcal{B}_{\eta} \geq -\kappa \mathcal{B}_{0}$, this provides little guidance as to what to do with h at the transition point. One possibility is to accommodate the transition through the use of the cloud fraction parameterization in the mass flux. Deep within the trade-wind regimes the parameterization proposed by Neggers et al., [27] acts simply to keep h near the η and for all practical purposes one might be able to model it as simply

$$\mathcal{A} = \operatorname{Erf}((h - \eta)/\sigma_h) \tag{68}$$

where σ_h is a prescribed length-scale (which could be a function of state) and Erf denotes the error function. In this model \mathcal{A} increases as h/σ_h becomes large relative to η/σ_h , thereby helping to keep h near η , by venting the ABL if h becomes too large. In this context further limiting vented cloud fraction as a function of the entrainment coefficient (α) could help provide a smooth and physically plausible transition of h.

As another alternative, we note that when $\Delta(s + L_v q) < 0$ vented cloud water can efficiently cool the air above the ABL thus promoting the development and growth of the trade-wind layer through the injection of liquid water by a venting cloud mass [33]. Roughly speaking $\Delta(s + L_v q) < 0$ transitions from modest positive values over the cold subtropical oceans, to large negative values as one moves over warmer waters. In this respect one could also explore using this (the moist static energy) limit as a basis for further modulating the ventilating cloud fraction in bulk models of the sub-cloud layer as a means of enforcing continuity between the trade-wind and stratocumulus regimes. A variant of this approach, often associated with the idea of cloud-top entrainment instability [30,12], is already used in one class of general circulation models and appears effective on climatological scales irrespective of whether or not such conditions are associated with the dessication of the stratocumulus layer on the timescales of large-eddies in the ABL.

6.1 Deep convection

As discussed above, the principal motivation for using of the sub-cloud layer as a basis for a bulk description of the boundary layer is to facilitate coupling to nearby regions of moist convection. Indeed, a variant of the sub-cloud model discussed in §5.2 is the basis for several descriptions of deep convection [13,32]. In the deep convective limit the only change to the equations is that precipitation, ΔF_q , is no longer negligible and the mass flux description:

$$M\Delta_{+}\phi = -\overline{w'\phi'}_{+} \tag{69}$$

ceases to be adequate as it fails to account for convective downdrafts. These downdrafts act to transport air with free tropospheric properties into the ABL, acting to homogenize the two layers. They are often incorporated by a two draft model of convective activity, such that

$$\overline{w'\phi'}_{+} = m_u(\widehat{\phi} - \phi_+) - m_d(\phi_f - \phi_+), \tag{70}$$

where $m_u > 0$ is an updraft mass flux and $m_d > 0$ is a downdraft mass flux, so that the net mass flux out of the ABL is $M = m_u - m_d$. If, following Raymond [32] we assume $m_d = \alpha_m m_u$ with $0 \le \alpha < 1$, and use the cloud model (56) then

$$-\overline{w'\phi'}_{+} = M\Delta_{+}\phi + M\frac{\alpha_m}{1 - \alpha_m}(\phi_f - \widehat{\phi}). \tag{71}$$

The second term on the rhs amounts to an additional source term in the ϕ equations. It acts to relax $\widehat{\phi}$ to ϕ_f on the timescale $h/m_d = h(1-\alpha_m)/(\alpha_m M)$ and, as expected, vanishes for zero downdraft mass flux $(\alpha_m = 0)$. In equilibrium the efficacy of this term is less substantial than one might imagine at a first glance, because an increased downdraft mass flux corresponds to less entrainment one can think of the fraction m_d/E of the entrainment flux $E(\widehat{\phi} - \phi_+)$ being replaced by a flux of $E(\widehat{\phi} - \phi_f)$. These will differ insofar as $\phi_f \neq \phi$. In the context of our mixing line model of the cloud layer, this difference is measured by β_+ , and matters only in so far as downdrafts are active, i.e., $\alpha_m > 0$. The effects of the downdrafts are perhaps best illustrated from the nature of the equilibrium solutions themselves, which substituting from above are:

$$h_{\infty} = \frac{E - M}{\mathcal{D}},\tag{72}$$

$$\widehat{s}_{\infty} = \frac{V s_0 + \widetilde{E} s_f - \Delta F_s}{V + \widetilde{E}}, \quad \text{and}$$
(73)

$$\widehat{q}_{\infty} = \frac{Vq_0 + \widetilde{E}q_f - h\widehat{\mathbf{u}} \cdot \nabla q}{V + \widetilde{E}e}.$$
(74)

The key difference between the above equilibria, and those for solutions (e.g., 61-62) without downdrafts is that in the thermodynamic budgets β_+E has been replaced by an effective entrainment velocity defined as:

$$\tilde{E} \equiv \beta_{+}E + M \frac{\alpha_{m}}{1 - \alpha_{m}}.$$
(75)

Because in convective regions $E_{\infty} \approx M_{\infty}$ this amounts to enhancing entrainment in this equilibrium by the factor $1 + \alpha_m \left(\beta_+^{-1} - 1\right)$ Thus as $\beta_+ \to 1$ the effect of the downdraft term goes to zero. Physically this is because downdrafts act to bring air with the property ϕ_f , rather than ϕ_+ , into the sub-cloud layer and for $\beta_+ = 1$, $\phi_+ = \phi_f$. In this sense α measures the degree to which the cloud-layer buffers the sub-cloud layer from the free troposphere, thereby allowing the two layers to decouple and differentiate themselves from one another. In the deep convective limit, when α_m is large the sub-cloud layer very effectively feels the properties of the free troposphere.

Some points worth noting from the above analysis are: (i) Equilibria are permitted in regions where $\mathcal{D} < 0$, they simply require that M > E; (ii) the solutions above are naturally continuous with those for regions of shallow convection, with α_m essentially measuring the amount of deep convection; (iii) although the focus on the sub-cloud layer for regions of shallow or deep convection would seem to avoid the problem of modeling the cloud layer, the cloud model is effectively specified by the parameters β_+ , α_m and the height z_f chosen for the determination of s_f . Moreover, it is reasonable to expect these parameters not to be independent of one another. For instance as the cloud layer and the free-troposphere become less differentiated deeper convection may ensue and β_+ may be expected to increase (so as to preserve a constant $\Delta_+ b$ at cloud base, along with α_m and z_f .

6.2 Momentum considerations

Until now we have limited our discussion exclusively to the thermodynamic state of the ABL. Its dynamic state, as represented by $\hat{\mathbf{u}}$ is also interesting. Not only is it fundamental in determining the surface fluxes, and ocean coupling, but it also determines areas of convergence and divergence, which in turn helps determine the types of ABL equilibria one may expect [19]. The concepts discussed above are, however, also very relevant to a treatment of low-level momentum, for which the bulk equations become:

$$h\frac{D\widehat{\mathbf{u}}}{Dt} - \Delta_{+}\mathbf{u} \left[\frac{Dh}{Dt} + \mathcal{D}h \right] + f\mathbf{k} \times \widehat{\mathbf{u}} = -\frac{1}{\rho_{0}} \left[\nabla \widehat{p} - p_{+} \nabla \ln h \right] - \overline{w'} \mathbf{u'}_{+} + V \Delta_{0} \widehat{\mathbf{u}}, \tag{76}$$

where f is an inverse timescale which meters the Coriolis force, \mathbf{k} is the unit vector pointing upwards, and p is the pressure. Neglecting the $\ln h$ term on the rhs and modeling the momentum flux out of the ABL as if momentum were a passive scalar (and hence well represented by the mass flux approach described above) (76) reduces to

$$h\frac{D\widehat{\mathbf{u}}}{Dt} + f\mathbf{k} \times \widehat{\mathbf{u}} = -\frac{1}{\rho_0} \nabla \widehat{p} + E\Delta_+ \widehat{\mathbf{u}} + V\Delta_0 \widehat{\mathbf{u}}, \tag{77}$$

which is similar to the bulk equation for thermodynamic quantities, except for the appearance of the Coriolis term and the pressure gradient (which depends on the state of the layer) as a source term. The ability of steady-states of this equation, for a fixed h=500 m and E=0.01 m s⁻¹ were shown to be a reasonable description of the relationship amongst surface pressure, winds and the geographic variation of f [43]. More systematic applications of this approach have also proven useful in describing the pattern of winds over the eastern tropical Pacific [23]. On a more quantitative level it remains unclear as to the extent to which

$$E\Delta_{+}\mathbf{u} = \Delta_{+}\mathbf{u} \left[\frac{Dh}{Dt} + \mathcal{D}h + M \right] \approx \Delta_{+}\mathbf{u} \left[\frac{Dh}{Dt} + \mathcal{D}h \right] - \overline{w'\mathbf{u'}}_{+}. \tag{78}$$

which presupposes that the last term on the rhs behaves as a mass flux following Eq.(10), e.g., $M\Delta_+\mathbf{u} = -\overline{w'\mathbf{u'}}_+$. Because $\mathbf{u'}/\widehat{\mathbf{u}}$ is much closer to unity than either s'/\widehat{s} or q'/\widehat{q} the ability of the circulation to organize fluctuations of \mathbf{u} in correspondence to the convection, even for regions of shallow convection, may critically determine the net momentum transport through the top of the ABL. These issues become even more acute for deep convection for which momentum fluctuations are on the order of the mean wind and both are organized by, and help organize, areas of enhanced convection.

6.3 Missing regimes

Our description of bulk representations of atmospheric ABLs is reasonably complete for tropical maritime regions, where the flow tends to be toward warmer water, so that the surface buoyancy flux, \mathcal{B}_0 , is positive and the sub-cloud layer is convective. However some important regimes are less obviously well described by the above approach. These include regions of stable stratification, such as in the equatorial cold-tongue region of the eastern Pacific, and over land at night. In regions of stable stratification, the ABL interacts in less predictable ways with regions of turbulence production aloft (such as jets, breaking gravity waves, etc.) and this formulation in which h is related to the generation of a wall bounded turbulent flow is less obviously relevant. Another problematic regime is over land, here the lack of available moisture and low surface conductivities can promote the development of very deep daytime cloud layers for which many of the advantages of the treatment of just the sub-cloud layer may be lost.

7 Summary & Conclusions

We have reviewed the representation of the atmospheric boundary layer (ABL) using bulk theory, focusing on the representation of the thermodynamic state of tropical and subtropical maritime regimes, away from regions of deep convection. Three paradigms were explored in some detail: (i) stratocumulus topped mixed layers as embodied by the theory first developed by Lilly [18]; (ii) the trade-wind boundary layer as represented by the model of Betts and Ridgway [6], and models of the sub-cloud layer such as initially explored by Betts [5] and more recently (in this volume) by Neggers $et\ al.$, [27]. For all three approaches we show that closure of the bulk budgets of mass, enthalpy and moisture effectively requires the specification of three exchange velocities V, M, and E defining surface exchange, mass flux out of the ABL and entrainment at the ABL top respectively. To the extent the state of the surface layer can be related to the state of the bulk layer as a whole V follows relatively straightforwardly from surface layer similarity theory. For stratocumulus-topped ABL and the Betts and Ridgway model of the trade-wind layer, the ABL top is identified with the top of the cloud layer, and hence M is set to zero. Thus all that remains is to specify E. For the Betts and Ridgway model this is accomplished by assuming a fixed boundary layer structure and constraining the cloud base fluxes. For the stratocumulus

topped ABL determinations of E also rely on assumptions about boundary layer structure (typically that it is well mixed) and are an area of active research. Recent proposals however are shown to work reasonably well and when coupled to the bulk theory yield equilibrium solutions which are plausible representations of the observed climatology. Moreover, much of this climatology can be rationalized by yet simpler closures which yield analytic solutions to the bulk equations as a function of a prescribed entrainment efficiencies.

Closure of bulk models of just the sub-cloud layer typically involve additional assumptions. Although the cloud layer is not meant to be modeled, its structure must still be assumed so as to relate the state of the atmosphere just above cloud base to its structure above cloud top. Moreover, in addition to modeling E, a model of the mass flux, M out of the layer must also be specified. Typically Eis modeled by assuming that cumulus clouds have a small area and do not interact strongly with the turbulence in the sub-cloud layer. In which case E is given based on similarity relationships developed for dry convective boundary layers growing into layers of uniform stratification. We show that models of M which principally act to maintain the depth of the ABL at the lifting condensation level produce reasonable equilibria, which are characterized by a sub-cloud layer whose depth does not, but whose mass fluxes do depend on how the cloud layer is modeled. A proposal for avoiding this unsatisfactory parameter dependence is to model the buoyancy jump at the top of the sub-cloud layer using relationships valid for the growth of dry convective boundary layers. So doing effectively constrains one's model of the cloud layer and results in explicit predictions of the cloud core mass fraction, namely that this fraction \mathcal{A} scales as $(2\kappa+1)/\mathrm{Ri}$ where κ is the entrainment flux ratio (typically 0.2) and Ri is the bulk Richardson number defined as $(\eta N/w_*)^2$ where η is the height of cloud base, N is the Brunt-Väisälä frequency in the cloud layer and w_* is the Deardorff velocity scale. For typical values of $Ri \approx 60$ this yields cloud core fractions of around 2\%, similar to those produced by large-eddy simulation.

Means of blending, or matching, the various bulk regimes are also briefly explored. Here we find that the equilibrium solutions of the Betts and Ridgway model correspond to typical entrainment efficiencies which are not incommensurate with those indicative of decoupling for equilibrium solutions of the bulk equations for stratocumulus mixed layer. As a result one fruitful approach to matching these theories may be model the assumed boundary layer structure as a function of a decoupling parameter, such as the ratio of the cloud base fluxes to the surface fluxes assuming for this purpose a stratocumulus regime. Likewise, to couple the bulk representation of stratocumulus topped boundary layers to models of the sub-cloud structure our analysis suggests that relating the mass flux out of the ABL to the decoupling parameter (so that it becomes zero in regions where stratocumulus are expected) might be an effective approach. This latter type of model is also shown to couple naturally to regions of deep convection, and to previous representations of the bulk winds in the ABL. As such it appears to provide the most promising basis for the incorporation of boundary layer concepts into simplified models of tropical dynamics.

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