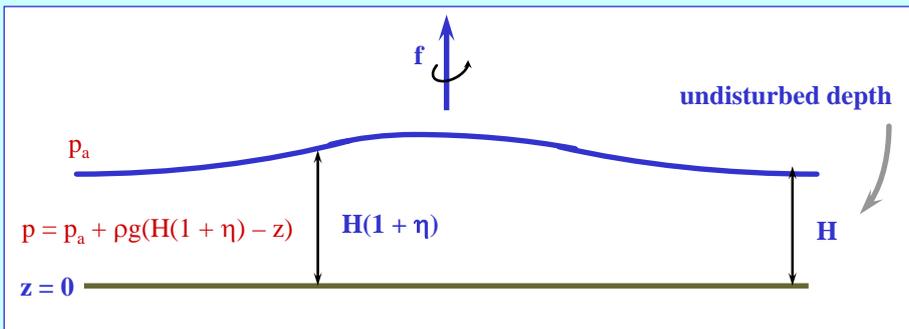


Reminder: Inertia-gravity waves

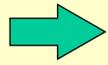


Shallow water model configuration

- In **pure inertial wave motion**, horizontal pressure gradients are zero.
- Consider now **waves** in a layer of rotating fluid with a **free surface** where horizontal pressure gradients are associated with free surface displacements.

Consider hydrostatic motions - then

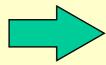
$$p(x, y, z, t) = p_a + \rho g [H\{1 + \eta(x, y, t)\} - z]$$



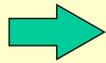
$$\frac{1}{\rho} \nabla_h p = gH \nabla_h \eta$$



independent of z



The fluid acceleration is independent of z .



**If the velocities are initially independent of z ,
then they will remain so.**

Linearized equations - no basic flow

$$\frac{\partial u}{\partial t} - fv = -gH \frac{\partial \eta}{\partial x}$$

$$\frac{\partial v}{\partial t} + fu = -gH \frac{\partial \eta}{\partial y}$$

$$\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} = -\frac{\partial \eta}{\partial t}$$

Consider wave motions which are independent of y .

A solution exists of the form $u = \hat{u} \cos(kx - \omega t)$,

$v = \hat{v} \sin(kx - \omega t)$, if 

$\eta = \hat{\eta} \cos(kx - \omega t)$,

\hat{u} , \hat{v} and $\hat{\eta}$ constants

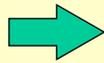
$$\omega \hat{u} - f \hat{v} - gHk \hat{\eta} = 0,$$

$$f \hat{u} - \omega \hat{v} = 0,$$

$$-k \hat{u} + \omega \hat{\eta} = 0,$$

These algebraic equations for \hat{u} , \hat{v} and $\hat{\eta}$ have solutions only if

$$\begin{vmatrix} \omega & -f & -gHk \\ f & -\omega & 0 \\ -k & 0 & \omega \end{vmatrix} = -\omega^3 + \omega(f^2 + gHk^2) = 0$$



$$\omega = 0 \text{ or } \omega^2 = f^2 + gHk^2$$

The solution with $\omega = 0$ corresponds with the steady solution ($\partial/\partial t = 0$) of the equations and represents a **steady current in strict geostrophic balance** in which

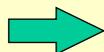
$$v = \frac{gH}{f} \frac{\partial \eta}{\partial x}$$

The other two solutions correspond with so-called **inertia-gravity waves**, with the dispersion relation

$$\omega^2 = f^2 + gHk^2$$

The phase speed of these is

$$c_p = \omega / k = \pm \sqrt{[gH + f^2 / k^2]}$$



The waves are **dispersive**

Inclusion of the beta effect

- Replace the v-momentum equation by the vorticity equation
=>

$$\frac{\partial u}{\partial t} - f_0 v = -gH \frac{\partial \eta}{\partial x}$$

$$\frac{\partial \zeta}{\partial t} + \beta v = f_0 \frac{\partial \eta}{\partial t}$$

$$\frac{\partial \eta}{\partial t} + \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} = 0$$

where

$$\zeta = \frac{\partial v}{\partial x} - \frac{\partial u}{\partial y}$$

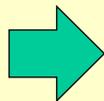
Assume that f can be approximated by its value f_0 at a particular latitude, **except** when differentiated with respect to y in the vorticity equation. This is justified provided that meridional particle displacements are small.

- Again assume that $\partial/\partial y \equiv 0$ and consider travelling-wave solutions of the form:

$$u = \hat{u} \cos(kx - \omega t),$$

$$v = \hat{v} \sin(kx - \omega t),$$

$$\eta = \hat{\eta} \cos(kx - \omega t),$$

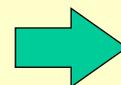


$$\omega \hat{u} - f_0 \hat{v} - gHk \hat{\eta} = 0,$$

$$(k\omega + \beta) \hat{v} - f_0 \omega \hat{\eta} = 0,$$

$$-k \hat{u} + \omega \hat{\eta} = 0,$$

These are consistent **only if the determinant is zero**



Now

expanding by the second row

$$\begin{vmatrix} \omega & -f_0 & -gHk \\ 0 & +(k\omega + \beta) & -\omega f_0 \\ -k & 0 & \omega \end{vmatrix} = (\omega^2 - gHk^2)(\omega k + \beta) - f_0^2 \omega k = 0$$

A cubic equation for ω with three real roots.

When $\omega \gg \beta/k$, the **two non-zero roots** are given approximately by the formula

$$\omega^2 = gHk^2 + f_0^2$$

This is precisely the **dispersion relation for inertia-gravity waves**.

When $\omega^2 \ll gHk^2$, there is **one root** given approximately by

$$\omega = -\beta k / (k^2 + f_0^2 / gH)$$

Filtering

Equations: ~~$\frac{\partial u}{\partial t} - f_0 v = -gH \frac{\partial \eta}{\partial x}$~~

$$\frac{\partial \zeta}{\partial t} + \beta v = f_0 \frac{\partial \eta}{\partial x}$$
$$\frac{\partial \eta}{\partial t} + \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} = 0$$
$$v = \frac{gH}{f_0} \frac{\partial \eta}{\partial x}$$

If $\partial/\partial y \equiv 0$ as before, the vorticity equation reduces to an equation for η , and since $\zeta = \partial v/\partial x \Rightarrow$

$$\frac{\partial}{\partial t} \left(\frac{gH}{f_0} \frac{\partial^2 \eta}{\partial x^2} - f_0 \eta \right) + \beta \frac{gH}{f_0} \frac{\partial \eta}{\partial x} = 0$$

This has the solution $\eta = \hat{\eta} \cos(kx - \omega t)$, where

Dispersion relation for a divergent planetary wave

$$\omega = -\beta k / (k^2 + f_0^2 / gH)$$

- There is no other solution for ω as there was before.
- In other words, making the geostrophic approximation when calculating v has **filtered** out in the inertia-gravity wave modes from the equation set, leaving only the low frequency planetary wave mode.
- This is not too surprising since the inertia-gravity waves, by their very essence, are not geostrophically-balanced motions.

Filtered equations more generally

- The idea of filtering sets of equations is an important one in geophysical applications.
- The **quasi-geostrophic equations** are often referred to as '**filtered equations**' since, as in the above analysis, the consequence of computing the horizontal velocity geostrophically from the pressure or stream-function suppresses the high frequency inertia-gravity waves which would otherwise be supported by the Boussinesq equations.
- Furthermore, the **Boussinesq equations** themselves form a **filtered system** in the sense that the approximations which lead to them filter out **compressible, or acoustic waves**.

The Balance Equations

- The full nonlinear form of the shallow-water equations in a layer of fluid of variable depth $h(x,y,t)$ is

$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} - fv = -g \frac{\partial h}{\partial x}$$

$$\frac{\partial v}{\partial t} + u \frac{\partial v}{\partial x} + v \frac{\partial v}{\partial y} + fu = -g \frac{\partial h}{\partial y}$$

$$\frac{\partial h}{\partial t} + u \frac{\partial h}{\partial x} + v \frac{\partial h}{\partial y} + h \left(\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} \right) = 0$$

$$\zeta = \partial_x v - \partial_y u \quad D = \partial_x u + \partial_y v$$

relative vorticity **horizontal divergence**

- The u- and v-equations can be replaced by the vorticity and divergence equations: =>

vorticity equation

$$\frac{\partial \zeta}{\partial t} + u \frac{\partial \zeta}{\partial x} + v \frac{\partial \zeta}{\partial y} + \beta v + (\zeta + f) D = 0$$

divergence equation

$$\frac{\partial D}{\partial t} + u \frac{\partial D}{\partial x} + v \frac{\partial D}{\partial y} + D^2 - f\zeta - 2J(u, v) + \beta u + g\nabla^2 h = 0,$$

$\beta = df/dy$

Combine the vorticity equation and the continuity equation to form the **potential vorticity equation**:

$$\frac{\partial q}{\partial t} + u \frac{\partial q}{\partial x} + v \frac{\partial q}{\partial y} = 0,$$

potential vorticity is $q = (\zeta + f)/h$

$$\frac{\partial \zeta}{\partial t} + u \frac{\partial \zeta}{\partial x} + v \frac{\partial \zeta}{\partial y} + \beta v + (\zeta + f) D = 0$$

$$\frac{\partial D}{\partial t} + u \frac{\partial D}{\partial x} + v \frac{\partial D}{\partial y} + D^2 - f\zeta - 2J(u, v) + \beta u + g\nabla^2 h = 0$$

$$\frac{\partial q}{\partial t} + u \frac{\partial q}{\partial x} + v \frac{\partial q}{\partial y} = 0$$

Equivalent to:

$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} - fv = -g \frac{\partial h}{\partial x}$$

$$\frac{\partial v}{\partial t} + u \frac{\partial v}{\partial x} + v \frac{\partial v}{\partial y} + fu = -g \frac{\partial h}{\partial y}$$

$$\frac{\partial h}{\partial t} + u \frac{\partial h}{\partial x} + v \frac{\partial h}{\partial y} + h \left(\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} \right) = 0$$

Scale analysis

➤ **Choose representative scales:**

U for the horizontal velocity components u, v ,

L for the horizontal length scale of the motion,

H for the undisturbed fluid depth,

δH for depth departures $h - H$,

f_0 for the Coriolis parameter, and

$T = L/U$ - an advective time scale.

➤ **Define two nondimensional parameters:**

the Rossby number, $Ro = U/(f_0 L)$, and

the Froude number, $Fr = U^2/(gH)$.

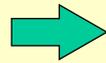
Treatment of f

Let $f = f_0(1 + \text{Ro}\beta'y)$,

where y is nondimensional and $\beta' = \beta L^2 / U$.

For middle latitude systems

$$U \sim 10 \text{ m s}^{-1}, \beta \sim 10^{-11} \text{ m}^{-1} \text{ s}^{-1}, L \sim 10^6 \text{ m}$$



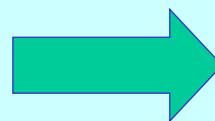
β' is of order unity.

The nondimensional forms of the u -momentum and continuity equations are:

$$\text{Ro} \left(\frac{\partial u}{\partial t} + \dots - \beta' y v \right) - v = - \frac{g \delta H}{f_0 U L} \frac{\partial h}{\partial x},$$

$$\frac{\delta H}{H} \left(\frac{\partial h}{\partial t} + \dots \right) + D = 0,$$

➤ Let us examine first the **quasi-geostrophic** scaling:



$$\text{Ro} [\cancel{\partial_t u + \dots - \beta' y v}] - v = -(g \delta H / f_0 UL) \partial_x h,$$

$$(\delta H / H) [\partial_t h + \dots] + D = 0,$$

- **In quasi-geostrophic motion** ($\text{Ro} \ll 1$), **the term proportional to Ro can be neglected.**
- **Then** $g\delta H/f_0 UL \approx O(1)$.
- **We could choose the scale H so that this quantity is exactly unity, i.e. $\delta H = f_0 UL/g$. Then the term $\delta H/H$ may be written**

$$\delta H / H = \text{Ro}^{-1} \text{Fr}.$$

The momentum equations are

$$\text{Ro} \left(\frac{\partial u}{\partial t} + \dots - \beta' y v \right) - v = - \frac{g \delta H}{f_0 UL} \frac{\partial h}{\partial x}$$

$$\text{Ro} \left(\frac{\partial v}{\partial t} + \dots - \beta' y v \right) + u = - \frac{g \delta H}{f_0 UL} \frac{\partial h}{\partial y}$$

⇒ **to lowest order in Ro**

$$v = \frac{\partial h}{\partial x} \quad u = - \frac{\partial h}{\partial y}$$



$D = 0$ (or, more generally, $D \leq O(\text{Ro})$).

$$\frac{\delta H}{H} \left(\frac{\partial h}{\partial t} + \dots \right) + D = 0,$$



a consistent scaling for quasi-geostrophic motion is that $\delta H/H = \text{Ro}^{-1} \text{Fr} = O(\text{Ro})$ so that the term proportional to $\delta H/H \rightarrow 0$ as $\text{Ro} \rightarrow 0$.

The vorticity and divergence equations

➤ In nondimensional form, the vorticity and divergence equations are

$$\text{Ro} \left(\frac{\partial \zeta}{\partial t} + \dots + \beta' v \right) + D + \text{Ro}(\beta' y + \zeta) D = 0$$

and

$$\text{Ro} \left(\frac{\partial D}{\partial t} + \dots \right) + D^2 - 2J(u, v) - \zeta(1 + \beta' \text{Ro} y) + \text{Ro} \beta' u + \nabla^2 h = 0,$$

where $J(u, v)$ is the **Jacobian** $\frac{\partial u}{\partial x} \frac{\partial v}{\partial y} - \frac{\partial v}{\partial x} \frac{\partial u}{\partial y}$

The quasi-geostrophic approximation

$$\text{Ro} \left(\frac{\partial \zeta}{\partial t} + \dots + \beta' v \right) + D + \text{Ro}(\beta' y + \zeta) D = 0$$

$$\text{Ro} \left(\frac{\partial D}{\partial t} + \dots + D^2 - 2J(u, v) \right) - \zeta(1 + \beta' \text{Ro} y) + \text{Ro} \beta' u + \nabla^2 h = 0$$

At lowest order in Rossby number

$$\text{Ro} \left(\frac{\partial \zeta}{\partial t} + \dots + \beta' v + D_1 \right) = 0 \quad \text{where } D = \text{Ro} D_1$$

$$\nabla^2 h = \zeta$$

When $\delta H/H = \text{Ro}^{-1} \text{Fr}$, the continuity equation \Rightarrow

$$\frac{\partial h}{\partial t} + \dots + \frac{1}{\mu} D_1 = 0 \quad \text{where } \mu = \text{Ro}^{-2} \text{Fr} \text{ and } D = \text{Ro} D_1.$$

$$\text{Ro} \left(\frac{\partial \zeta}{\partial t} + \dots + \beta' v + D_1 \right) = 0$$

$$\nabla^2 h = \zeta$$

$$\frac{\partial h}{\partial t} + \dots - \frac{1}{\mu} D_1 = 0$$

- These are the **quasi-geostrophic forms of the vorticity, divergence and continuity equations.**
- Note that the divergence equation has reduced to a diagnostic one relating the fluid depth to the vorticity which is consistent with ζ obtained from

$$\frac{\partial \zeta}{\partial t} + u \frac{\partial \zeta}{\partial x} + v \frac{\partial \zeta}{\partial y} + \beta v + (\zeta + f) D = 0$$

- $D = 0$ at $O(\text{Ro}^0) \Rightarrow$ there exists a **streamfunction** ψ such that

$$u = -\frac{\partial \psi}{\partial y} \quad v = \frac{\partial \psi}{\partial x}$$

and from

$$\text{Ro} \left(\frac{\partial u}{\partial t} + \dots - \beta' y v \right) - v = -\frac{g \delta H}{f_0 U L} \frac{\partial h}{\partial x}$$

$$\psi = h + \text{constant (e.g. } h - H).$$

The potential vorticity equation

- The **potential vorticity equation** in nondimensional form can be obtained by eliminating D_1 between

$$\text{Ro} \left(\frac{\partial \zeta}{\partial t} + \dots + \beta' v + D_1 \right) = 0 \quad \text{and} \quad \frac{\partial h}{\partial t} + \dots + \frac{1}{\mu} D_1 = 0$$

➔ a single equation for ψ (or h) =>

$$\left(\frac{\partial}{\partial t} + \mathbf{u} \frac{\partial}{\partial x} + \mathbf{v} \frac{\partial}{\partial y} \right) (\nabla^2 \psi + \beta' y - \mu \psi) = 0$$

analogous to the form for a stably-stratified fluid.

Note: in the case with a continuous vertical stratification, the ψ term would be replaced by a second-order vertical derivative of ψ .

- The **quasi-geostrophic approximation** leads to an elegant mathematical theory, but calculations based upon it tend to be inaccurate in many atmospheric situations such as when the isobars are strongly curved.
- In the latter case, we know that centrifugal forces are important and gradient wind balance gives a more accurate approximation.
- We try to improve the quasi-geostrophic theory by including terms of **higher order** in Ro in the equations.

- We decompose the horizontal velocity into **rotational** and **divergent** components

$$\mathbf{u} = \mathbf{u}_\psi + \mathbf{u}_\chi$$

$$\mathbf{u}_\psi = \mathbf{k} \wedge \nabla \psi$$

$$\mathbf{u}_\chi = \nabla \chi$$

ψ is the **streamfunction**.

χ the **velocity potential**

It follows that

$$\zeta = \nabla^2 \psi \quad \text{and} \quad D = \nabla^2 \chi$$

- This decomposition is general (see Holton, 1972, Appendix), but is **not unique** : \Rightarrow
- One can add equal and opposite flows with zero vorticity and divergence to the two components without affecting the total velocity \mathbf{u} .
- Consistent with the quasi-geostrophic scaling, where χ is zero to $O(\text{Ro}^0)$, we scale χ with ULRo so that in nondimensional form

$$\mathbf{u} = \mathbf{u}_\psi + \mathbf{u}_\chi \quad \rightarrow \quad \mathbf{u} = \mathbf{u}_\psi + \text{Ro} \mathbf{u}_\chi$$

Equations $\text{Ro}[\partial_t \zeta + \dots + \beta'v] + D + \text{Ro}(\beta'y + \zeta)D = 0,$

$$\begin{aligned} &\text{Ro}[\partial_t D + \dots + D^2 - 2J(\mathbf{u}, \mathbf{v})] - \zeta(1 + \beta'\text{Ro}y) \\ &+ \text{Ro}\beta'u + \nabla^2 h = 0, \end{aligned}$$

$$\begin{aligned} \rightarrow &[\partial_t \nabla^2 \psi + (\mathbf{u}_\psi + \text{Ro} \cdot \mathbf{u}_\chi) \cdot \nabla(\nabla^2 \psi) + \beta' \partial_x \psi] + \nabla^2 \chi \\ &+ (\beta'y + \nabla^2 \psi) \nabla^2 \chi = 0, \end{aligned}$$

and

$$\begin{aligned} &\text{Ro}^2[\partial_t \nabla^2 \chi + J(\psi, \nabla^2 \chi)] + \text{Ro}^3[\mathbf{u}_\chi \cdot \nabla(\nabla^2 \chi) + (\nabla^2 \chi)^2] - \\ &2\text{Ro}J(\mathbf{u}_\psi, \mathbf{v}_\psi) - 2\text{Ro}^2[J(\mathbf{u}_\chi, \mathbf{v}_\psi) + J(\mathbf{u}_\psi, \mathbf{v}_\chi)] - 2\text{Ro}^3 J(\mathbf{u}_\chi, \mathbf{v}_\chi) - \\ &\nabla^2 \psi(1 + \beta'\text{Ro}y) + \beta'\text{Ro} \mathbf{u}_\psi + \beta'\text{Ro}^2 \mathbf{u}_\chi + \nabla^2 h = 0. \end{aligned}$$

The idea is to neglect terms of order Ro^2 and Ro^3 in these equations, together with the equivalent approximation in the continuity equation

$$\partial_t h + u \partial_x h + v \partial_y h + h(\partial_x u + \partial_y v) = 0,$$

In dimensional form they may be written

$$[\partial_t + \mathbf{u} \cdot \nabla](\nabla^2 \psi + f) + (\nabla^2 \psi + f) \nabla^2 \chi = 0$$

$$2[(\partial_{xx} \psi)(\partial_{yy} \psi) - (\partial_{xy} \psi)^2] + \nabla \cdot (f \nabla \psi) - \nabla^2 h = 0$$

$$\partial_t h + \mathbf{u} \cdot \nabla h + h \nabla^2 \chi = 0$$

$$\mathbf{u} = \mathbf{u}_\psi + \mathbf{u}_\chi.$$

Note: the divergence equation has been reduced to a diagnostic one relating ψ to h .

Moreover the advection of $\nabla^2 \psi + f$ in

$$[\partial_t + \mathbf{u} \cdot \nabla](\nabla^2 \psi + f) + (\nabla^2 \psi + f) \nabla^2 \chi = 0$$

and of h in $\partial_t h + \mathbf{u} \cdot \nabla h + h \nabla^2 \chi = 0$

is by the total wind \mathbf{u} and **not** just the nondivergent component of \mathbf{u} as in the quasi-geostrophic approximation.

The Balance Equations

$$\left(\frac{\partial}{\partial t} + \mathbf{u} \cdot \nabla \right) (\nabla^2 \psi + f) + (\nabla^2 \psi + f) \nabla^2 \chi = 0$$

$$2 \left(\frac{\partial^2 \psi}{\partial x^2} \frac{\partial^2 \psi}{\partial y^2} - \left(\frac{\partial^2 \psi}{\partial x \partial y} \right)^2 \right) + \nabla \cdot (f \nabla \psi) - g \nabla^2 h = 0$$

$$\frac{\partial h}{\partial t} + \mathbf{u} \cdot \nabla h + h \nabla^2 \chi = 0$$

- These equations are called **the balance equations**.
- They were first discussed by Charney (1955) and Bolin (1955).

The Balance Equations

- It can be shown that for a **steady axisymmetric flow on an f-plane, the equation:**

$$2 \left(\frac{\partial^2 \psi}{\partial x^2} \frac{\partial^2 \psi}{\partial y^2} - \left(\frac{\partial^2 \psi}{\partial x \partial y} \right)^2 \right) + \nabla \cdot (f \nabla \psi) - \nabla^2 h = 0$$

reduces to the gradient wind equation.

- => we may expect the equations to be a better approximation than the quasi-geostrophic system for **strongly curved flows**.

Methods of solution

- Unfortunately it is not possible to combine the balance equations into a single equation for ψ as in the quasi-geostrophic case and they are rather difficult to solve.
- Methods of solution are discussed by Gent and McWilliams (1983).
- **Note:** although the balanced equations were derived by truncating terms of $O(Ro^2)$ and higher, the only equation where approximation is made is the divergence equation.
- \Rightarrow the equations represent an approximate system valid essentially for **sufficiently small horizontal divergence**.
- As long as this is the case, the Rossby number is of no importance.

- Elimination of $\nabla^2\chi$ between

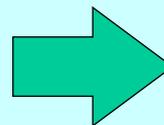
$$\frac{\partial h}{\partial t} + \mathbf{u} \cdot \nabla h + h \nabla^2 \chi = 0$$

and

$$\frac{\partial q}{\partial t} + u \frac{\partial q}{\partial x} + v \frac{\partial q}{\partial y} = 0$$

gives the **potential vorticity equation:**

Thus an alternative form of the balance system is



$$\left. \begin{aligned}
 q &= \frac{\zeta + f}{h} \\
 2 \left(\frac{\partial^2 \psi}{\partial x^2} \frac{\partial^2 \psi}{\partial y^2} - \left(\frac{\partial^2 \psi}{\partial x \partial y} \right)^2 \right) + \nabla \cdot (f \nabla \psi) - g \nabla^2 h &= 0
 \end{aligned} \right\} \rightarrow \psi \text{ and } h$$

$$\frac{\partial h}{\partial t} + \mathbf{u} \cdot \nabla h + h \nabla^2 \chi = 0$$

Given q , the first two equations can be regarded as a pair of simultaneous equations for diagnosing ψ and h , subject to appropriate boundary conditions.

Equation $\frac{\partial q}{\partial t} + u \frac{\partial q}{\partial x} + v \frac{\partial q}{\partial y} = 0$ enables the prediction of q ,

while $\frac{\partial h}{\partial t} + \mathbf{u} \cdot \nabla h + h \nabla^2 \chi = 0$ may be used to diagnose χ .

The Linear Balance Equations

Under certain circumstances [e.g. $Ro \ll 1$, $\beta' > O(1)$], the nonlinear terms in Eq.(11.47) may be neglected in which case the equation becomes

$$\nabla \cdot (f \nabla \psi) - g \nabla^2 h = 0$$

With this approximation, the system (11.46), (11.48) and (11.49) or (11.30), (11.31), (11.48), (11.49) constitute the **linear balance equations**.

These systems are considerably easier to solve than the balance equations.