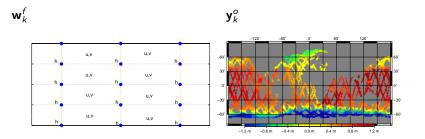
Accounting for small scale variability within data assimilation

Tijana Janjić

January 15, 2013

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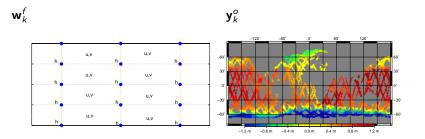
Data assimilation algorithm combine forecast and observations to produce the best analysis



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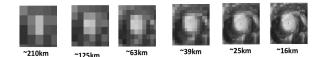
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Our goal: Best analysis for a prediction.

One major contributor to the forecast uncertainty is the model error.





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Model resolution (slide ECMWF)

Model error

Unfortunately model error statistics are not perfectly known and their determination remains a major challenge in assimilation systems.

Reasons behind the model error:

- accuracy of numerical schemes
- unrepresented subgrid scale processes
- inaccurate forcing and boundary conditions
- representation of orography as well as parametrisation uncertainty.

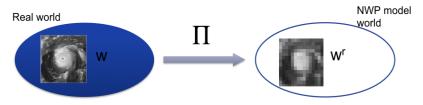
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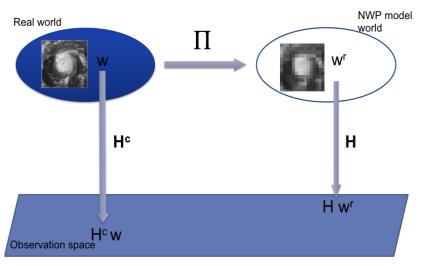
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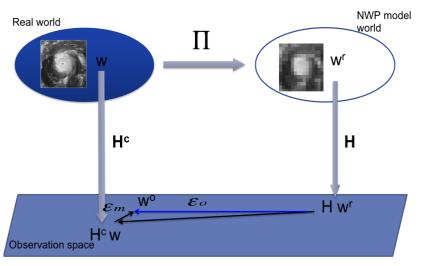
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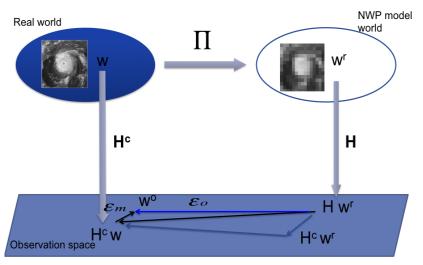
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Model error statistics produced by use of multiple physics packages, inclusion of stochastic kinetic energy backscatter scheme, parameter variations, as well as use of deterministic stochastic dynamical models (Berner et al. 2011).









Atmospheric Data Assimilation

The state $\mathbf{w} \equiv \mathbf{w}(\mathbf{x}, t)$ of the atmosphere at time t_k :

$$\mathbf{w}(\mathbf{x}, t_k) \equiv \begin{bmatrix} w_1(\mathbf{x}, t_k) \\ \vdots \\ w_q(\mathbf{x}, t_k) \end{bmatrix}$$

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where $w_i : \mathbb{R}^3 \times \mathbb{R} \to \mathbb{R}$, $\forall i = 1, ..., q$, $w_i \in \mathcal{B}$. (\mathcal{B} is a vector space of scalar valued, continuous functions.)

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Discrete problem: Find an estimate of some projection Πw of w on the space of the dynamical model.

$$\mathbf{\Pi}\mathbf{w}(\mathbf{x}, t_k) \equiv \begin{bmatrix} \Pi w_1(\mathbf{x}, t_k) \\ \vdots \\ \Pi w_q(\mathbf{x}, t_k) \end{bmatrix}$$

where $\Pi w_i : \mathbb{R}^3 \times \mathbb{R} \to \mathbb{R}, \forall i = 1, ..., q$. $\Pi w_i(\mathbf{x}, t_k) \in \mathcal{B}_N$, where \mathcal{B}_N is an *N*-dimensional subspace of \mathcal{B} .

 ϵ_k^o consists of measurement error and representativeness error. It can be divided into three parts:

$$\epsilon_k^o = \epsilon_k' + \epsilon_k'' + \epsilon_k^m$$

where

$$\begin{aligned} \epsilon'_k &\equiv & \mathsf{H}^c_k \mathsf{w}(\cdot, t_k) - \mathsf{H}^c_k \mathsf{\Pi} \mathsf{w}(\cdot, t_k) \\ &= & \mathsf{H}^c_k (\mathsf{I} - \mathsf{\Pi}) \mathsf{w}(\cdot, t_k) \end{aligned}$$

 ϵ'_k – will be called *error due to unresolved scales*.

$$\begin{aligned} \epsilon_k'' &\equiv \mathbf{H}_k^c \mathbf{\Pi} \mathbf{w}(\cdot, t_k) - \mathbf{H}_k \mathbf{\Pi} \mathbf{w}(\cdot, t_k) \\ &= [\mathbf{H}_k^c - \mathbf{H}_k] \mathbf{\Pi} \mathbf{w}(\cdot, t_k) \end{aligned}$$

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 ϵ_k'' – will be called *forward interpolation error*.

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- ▶ and it is state and time dependent (Janjic 2001, Janjic and Cohn 2006)

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- ► Example: for 40 × 40 km Radiosonde/Dropsonde wind observation, observational error < 0.5 m/s and Assigned error: 2 3 m/s</p>

2 Estimation of $\Pi w(\mathbf{x}, t)$

$$w(\mathbf{x},t) = w^{r}(\mathbf{x},t) + w^{u}(\mathbf{x},t),$$

with

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Dynamics:

$$w^{r}(\mathbf{x},t_{k+1})=F_{k+1,k}^{r}w^{r}(\mathbf{x},t_{k}),$$

Suppose $w^{u}(\mathbf{x}, t_{k})$ satisfies the dynamics

$$w^{u}(\mathbf{x}, t_{k+1}) = F^{u}_{k+1,k}w^{u}(\mathbf{x}, t_{k}) + F^{sl}_{k+1,k}w^{r}(\mathbf{x}, t_{k}).$$

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Observations: Assume $\mathbf{H}_k = \mathbf{H}_k^c |_{\mathcal{B}_N}$ (i.e., $\epsilon_k'' = 0$), $\mathbf{w}_k^o = \mathbf{H}_k^c w'(\cdot, t_k) + \epsilon_k^o$,

where ϵ_k^o is

$$egin{array}{rcl} \epsilon^o_k &=& \mathbf{H}^c_k w^u(\cdot,t_k)+\epsilon^m_k\ &=& \epsilon'_k+\epsilon^m_k. \end{array}$$

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A Kalman filter method

Augmented vector [Janjic 2001, Janjic and Cohn 2006]

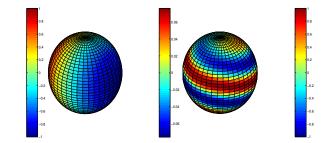
$$\mathbf{w}(\mathbf{x}, t_k) \equiv \begin{bmatrix} w^r(\mathbf{x}, t_k) \\ w^u(\mathbf{x}, t_k) \end{bmatrix},$$
$$\begin{bmatrix} w^r \\ w^u \end{bmatrix} (\mathbf{x}, t_{k+1}) = \begin{bmatrix} F^r_{k+1,k} & 0 \\ F^{ur}_{k+1,k} & F^u_{k+1,k} \end{bmatrix} \begin{bmatrix} w^r \\ w^u \end{bmatrix} (\mathbf{x}, t_k),$$

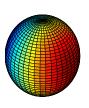
with observations given by

$$\mathbf{w}_k^o = egin{bmatrix} \mathbf{W}_k^c & \mathbf{H}_k^c \end{bmatrix} egin{bmatrix} w^r(m{\cdot},t_k) \ w^u(m{\cdot},t_k) \end{bmatrix} + m{\epsilon}_k^m.$$

This way the correct equations can be derived to take into account unresolved scales.

We would require estimates of correlation between resolved and unresolved scales as well as an estimates of the unresolved covariance at observation points.





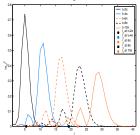
Initial state

Left: The state $w^r(\lambda, \varphi, t)$ which is being estimated. Right: The full state $w(\lambda, \varphi, t)$ from which the observations are taken.

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A method for estimation of time varying variance W^{uu} (Janjic 2001)

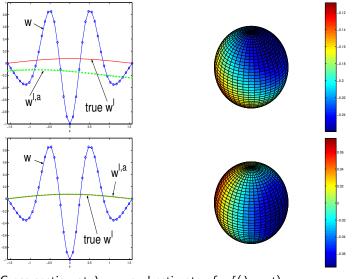


$$\sum_{l=l_c(t)-T_1}^{l_c(t)+T_2}\sum_{\rho=l_c(t)-T_1}^{l_c(t)+T_2}\sum_{m=-1}^1\sum_{q=-1}^1 \langle \widehat{w}_{lm}(t)\widehat{w}_{\rho q}(t)\rangle \ Y_l^m(\lambda,\varphi)Y_\rho^q(\lambda,\varphi).$$

The characteristic wavenumber will be calculated using the definition

$$I_{c}(t)^{2}(I_{c}(t)+1)^{2} \equiv \frac{\int_{0}^{2\pi} \int_{-\pi/2}^{\pi/2} \langle \Delta w^{u}(\lambda,\varphi,t) \Delta w^{u}(\lambda,\varphi,t) \rangle \cos \varphi \, d\varphi \, d\lambda}{\int_{0}^{2\pi} \int_{-\pi/2}^{\pi/2} \langle w^{u}(\lambda,\varphi,t) w^{u}(\lambda,\varphi,t) \rangle \cos \varphi \, d\varphi \, d\lambda}$$

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Cross section at $\lambda = \pi$ and estimate of $w^r(\lambda, \varphi, t)$. Upper: $\mathbf{H}_{2k}^c [\mathbf{H}_{1k}^c W_k^{uu}(\cdot, \cdot)]^T$ approximated by zero. Lower: $\mathbf{H}_{2k}^c [\mathbf{H}_{1k}^c W_k^{uu}(\cdot, \cdot)]^T$ approximated by adaptive method.

Estimating $\mathbf{H}_{2k}^{c}[\mathbf{H}_{1k}^{c}W_{k}^{uu}(\cdot,\cdot)]^{T}$

Desroziers et al., 2005:

- One source of information on the observation and background errors is contained in the statistics of the differences between observations and their background fields.
- This paper presents a set of diagnostics based on combinations of observation minus- background (OB), observation-minus-analysis (OA) and background-minus analysis (BA) differences.

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Problem that all deficiencies in the system are treated as one.

Estimating $\mathbf{H}_{2k}^{c}[\mathbf{H}_{1k}^{c}W_{k}^{uu}(\cdot,\cdot)]^{T}$

- Liu and Rabier et al. 2005: 1D idealized study for inferring statistics for satellite observations.
- Hamill and Whitaker 2005: Estimate the effect of unresolved scales by using sequence of lower model resolution then NWP model.

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Conclusion

- It was possible to get good results even with rather crude approximations of the unresolved scales.
- We require modeling of the covariance matrix obtained by evaluating this covariance function at the observation points.
- This covariance matrix depends on the properties of the observations and it is time and state dependent.
- Scalability of statistics is required.
- The error due to unresolved scales could not be neglected, or kept stationary.

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