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# LETKF for COSMO-DE: recent developments

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# Outline

- ▶ Multistep analysis
  - ▶ motivation and status
  - ▶ theory
- ▶ Status of KENDA
  - ▶ general setup
  - ▶ first results
  - ▶ finished and current experiments
  - ▶ radar operator: status & plans
- ▶ Outlook, open questions, discussion

## multistep analysis: motivation

also known as *successive, serial or batch assimilation*, but so far used for computational/algorithmic reasons.

For COSMO-LETKF various motivations (not completely independent) to use *multistep analysis*:

- ▶ local / nonlocal observations (e.g. Radiances, Christoph's idea)
- ▶ in relation with *adaptive localization*: different observation densities (conventional / radar)
- ▶ different observed scales (synoptic / convectional scale), observation errors

status: technically implemented / tested in COSMO-LETKF

- ▶ next step: test with radar data

paper: (together with África Perriáñez); prove equivalence of 1-step/multi-step for (ensemble) KF; investigate effect of localization

# multistep analysis: theory

## Theorem

*For the standard Kalman Filter with analysis  $\varphi^{(a)}$  at time  $t$ , and the multistep Kalman Filter with analysis  $\varphi^{(a,\xi)}$  for  $\xi = 1, \dots, q$  we have*

$$\varphi^{(a)} = \varphi^{(a,q)} \quad \text{and} \quad B^{(a)} = B^{(a,q)}.$$

## Theorem

*For the covariance matrices  $B^{(a)} := Q^{(a)}(Q^{(a)})'$  generated by the classical EnKF and the covariance matrix  $B^{(a,q)} := Q^{(a,q)}(Q^{(a,q)})'$  by the multi-step EnKF we have*

$$B^{(a)} = B^{(a,q)}.$$

# multistep analysis: theory

We define

$$A_1 := (Q^{(b)})'(H^{(1)})'(R^{(1)})^{-1}H^{(1)}Q^{(b)}$$

and

$$A_2 := (Q^{(b)})'(H^{(2)})'(R^{(2)})^{-1}H^{(2)}Q^{(b)}.$$

## Theorem (Multistep-EnKF Equivalence)

*Assume that the observation operators  $H^{(1)}$  and  $H^{(2)}$  for two different sets of measurements satisfy  $A_1A_2 = A_2A_1$ . Then the analysis ensemble generated by the multi-step EnKF with square root filter is identical to the analysis ensemble generated by the classical EnKF.*

# LETKF general setup

	<b>GME</b>	<b>COSMO</b>
ensemble member	40 + 1 (3dVar)	40 + 1 (det run)
horizontal resolution (ens)	ni128 ( $\approx$ 60 km)	2.8 km
horizontal resolution (det)	ni256 ( $\approx$ 30 km)	2.8 km
horiz. local. length scale	300 km	100 km
vert. local. length scale (ln p)	0.3 (0.075-0.5)	0.3 (0.075-0.5)
adapt. horiz. local.	not tested	tested (new exp)
additive model error	T (3dVar <b>B</b> )	F
(adaptive) inflation	T	T
conventional obs	T	T
Radiances	T (AMSU-A)	F
GPS-RO	new exps	F
Radar data	F	operator implemented
cloud height	F	Annika's talk
update frequency	3h	1h ( $\rightarrow$ 30/15 min)

# KENDA status: summary

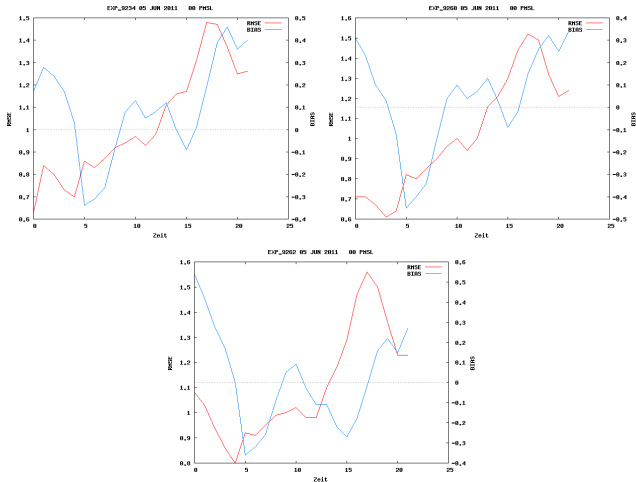
3 experiments so far:

- ▶ 9125 (base experiment)
- ▶ 9203 (modified observation errors)
- ▶ 9259 (saturation adjustment switched on, slightly modified observation errors, no assimilation of T2M, RH2M, (weak) adaptive localization used)

Verification:

- ▶ deterministic forecast: Klaus Stephan runs forecast up to 21 h, results comparable to nudging
- ▶ EPS: Richard Keane has run KENDA Ensemble from exp9259 for 2011060100 UTC; rmse larger than for COSMO-DE EPS, spread smaller

# RMSE/BIAS of deterministic KENDA forecasts

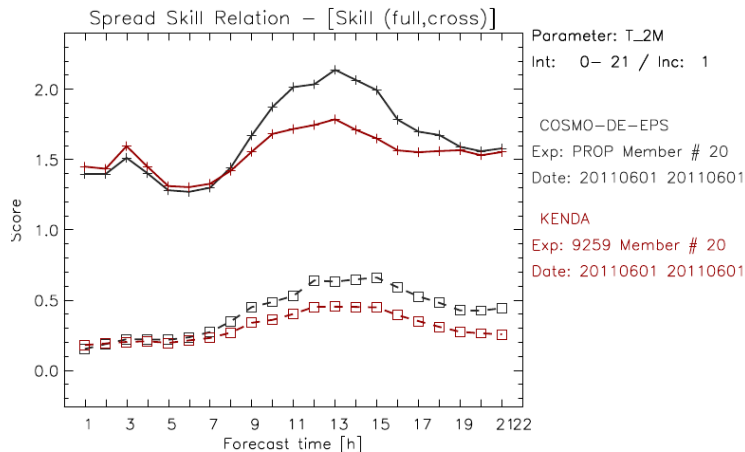


RMSE and BIAS of surface pressure, verified against SYNOP stations for LETKF, nudging and free forecast

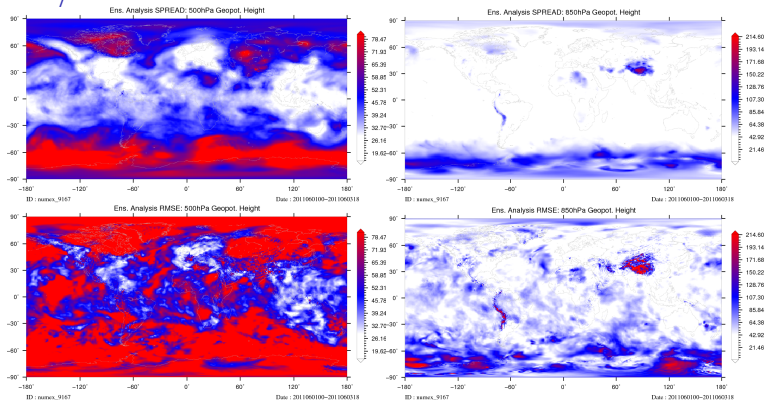
LETKF comparable to nudging



# KENDA EPS: comparison with COSMO-DE EPS



# RMSE/SPREAD of GME-LETKF

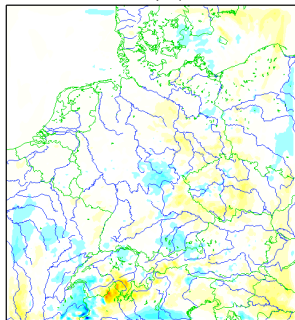


SPREAD and RMSE of GME-LETKF analysis (geop. height, 500 and 850 hPa)  
very low SPREAD over Europe and other data-rich areas → BC for  
COSMO-LETKF will also suffer from lack of spread; test/tune adaptive  
methods

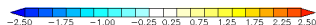
# KENDA: influence of boundary conditions I

Difference between:  
(9260: Nudging w/o LHN)–COSMO\_DE (9203: LETKF Det.  
initial: 01 JUN 2011 00 UTC  
valid: 01 JUN 2011 00 UTC

PMSL (hPa)

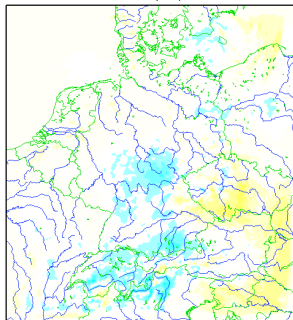


(1) Mean: 0.047014 Min: -2.90656 Max: 2.27562 Var: 0.0747623

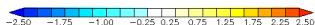


Difference between:  
(9298: noLHN Analyse)–COSMO\_DE (9259: LETKF Det. Anc  
initial: 01 JUN 2011 00 UTC  
valid: 01 JUN 2011 00 UTC

PMSL (hPa)



(1) Mean: 0.0440817 Min: -1.31 Max: 1.01313 Var: 0.0562013

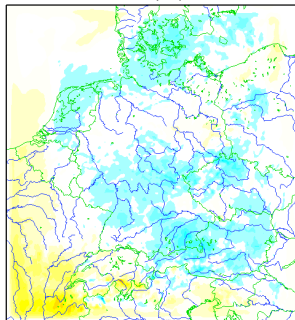


left: difference between nudging w/o LHN and LETKF with BC from COSMO-EU (nudging) and deterministic GME-LETKF (LETKF; currently 3dVar); right: same BC (GME-LETKF). (PMSL at 00 forecast time)

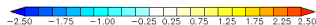
# KENDA: influence of boundary conditions II

Difference between:  
(9260: Nudging w/o LHN)–COSMO\_DE (9203: LETKF Det. Anc)  
initial: 01 JUN 2011 00 UTC  
valid: 01 JUN 2011 01 UTC

PMSL (hPa)

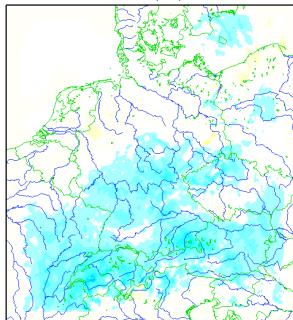


(1) Mean: -0.045613 Min: -1.83125 Max: 1.98188 Var: 0.128032

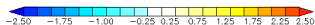


Difference between:  
(9298: noLHN Analyse)–COSMO\_DE (9259: LETKF Det. Anc)  
initial: 01 JUN 2011 00 UTC  
valid: 01 JUN 2011 01 UTC

PMSL (hPa)



(1) Mean: -0.175897 Min: -1.73375 Max: 0.777188 Var: 0.0666366

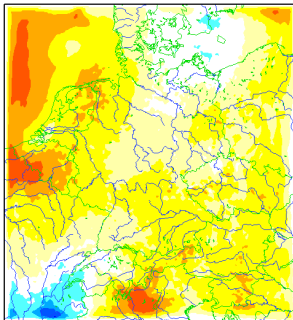


same as before, but for 01 forecast time

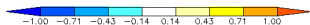
# KENDA: influence of boundary conditions III

Difference between:  
(9260: Nudging w/o LHN)–COSMO\_DE (9203: LETKF Det.  
initial: 01 JUN 2011 00 UTC  
valid: 01 JUN 2011 00 UTC

500hPa GPH

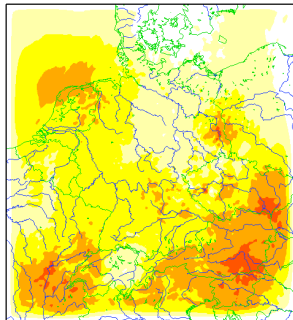


(1) Mean: 0.41607 Min: -1.05951 Max: 1.3994 Var: 0.0877838

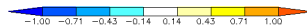


Difference between:  
(9298: noLHN Analyse)–COSMO\_DE (9259: LETKF Det. Anc  
initial: 01 JUN 2011 00 UTC  
valid: 01 JUN 2011 00 UTC

500hPa GPH



(1) Mean: 0.475635 Min: -0.160431 Max: 1.23515 Var: 0.0723127

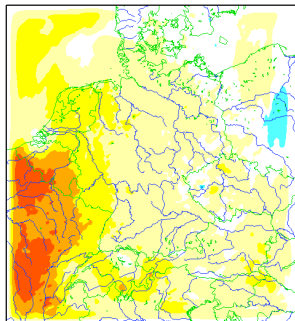


same as before, but for geopotential at 500 hPa, 00 forecast time

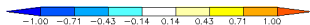
# KENDA: influence of boundary conditions IV

Difference between:  
(9260: Nudging w/o LHN)–COSMO\_DE (9203: LETKF Det.  
initial: 01 JUN 2011 00 UTC  
valid: 01 JUN 2011 01 UTC

500hPa GPH

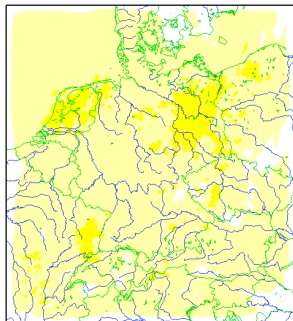


(1) Mean: 0.325534 Min: -0.691617 Max: 1.33215 Var: 0.9770969

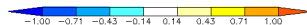


Difference between:  
(9298: noLHN Analyse)–COSMO\_DE (9259: LETKF Det. Anc  
initial: 01 JUN 2011 00 UTC  
valid: 01 JUN 2011 01 UTC

500hPa GPH



(1) Mean: 0.267564 Min: -0.226491 Max: 0.690303 Var: 0.0142421



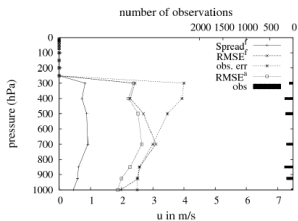
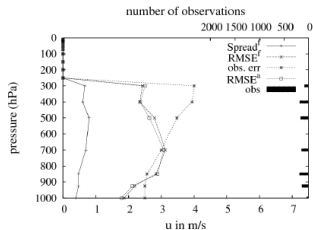
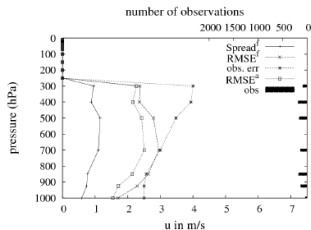
same as before, but for geopotential at 500 hPa, 01 forecast time

# radar operator

First results from Yuefei Zeng:

- ▶ radar data (radial wind) assimilated for 2011053118 UTC (1 analysis only, no cycling)
- ▶ 3 Experiments  $E0$ ,  $E1$ ,  $E2$ .
- ▶ all experiments use conventional data, settings are:
  - ▶  $E0$  radar passive
  - ▶  $E1$  radar active , localization length scale  $100km$  for conventional/radar
  - ▶  $E3$  radar active , localization length scale  $100km$  for conventional data,  $20km$  for radar
- ▶ no multistep analysis, but different localization radii used within 1 analysis

# radar operator



verification against AIREP for  $u$  wind component, Experiments E0, E1, E2



# Outlook

- ▶ *multistep analysis*: test with radar data (together with Yuefei Zeng), continue with theoretical work (paper), tests with toy models (more advanced models than Lorenz 95 needed)
- ▶ *technical (data base) problems* need to be solved to run experiments...; stand alone ( $\approx 1$  week) as alternative (store data in *ECFS*, use Thomas' stand-alone script)
- ▶ first results from KENDA (summary):
  - ▶ *deterministic*: in general comparable with nudging, but differences for surface pressure/geopotential (hydrostatic balancing, BC?)
  - ▶ *ensemble* not as good as COSMO-DE EPS, esp. lack of spread (due to BC / interior?); but only 1 forecast evaluated...
- ▶ *additional observations*: radar obs (radial winds, reflectivity), cloud height (Annika's talk)

# LETKF Theory

- ▶ do analysis in the  $k$ -dimensional ensemble space

$$\bar{\mathbf{w}}^a = \tilde{\mathbf{P}}^a (\mathbf{Y}^b)^T \mathbf{R}^{-1} (\mathbf{y} - \bar{\mathbf{y}}^b)$$
$$\tilde{\mathbf{P}}^a = [(k-1)\mathbf{I} + (\mathbf{Y}^b)^T \mathbf{R}^{-1} \mathbf{Y}^b]^{-1}$$

- ▶ in model space we have

$$\bar{\mathbf{x}}^a = \bar{\mathbf{x}}^b + \mathbf{X}^b \bar{\mathbf{w}}^a$$
$$\mathbf{P}^a = \mathbf{X}^b \tilde{\mathbf{P}}^a (\mathbf{X}^b)^T$$

- ▶ Now the analysis ensemble perturbations - with  $\mathbf{P}^a$  given above - are obtained via

$$\mathbf{X}^a = \mathbf{X}^b \mathbf{W}^a,$$

where  $\mathbf{W}^a = [(k-1)\tilde{\mathbf{P}}^a]^{1/2}$

# LETKF Theory

- ▶ it's possible to obtain a *deterministic run* via

$$\mathbf{x}_a^{det} = \mathbf{x}_b^{det} + \mathbf{K} \left[ \mathbf{y} - H(\mathbf{x}_b^{det}) \right]$$

with the *Kalman gain*  $\mathbf{K}$ :

$$\mathbf{K} = \mathbf{X}_b \left[ (k-1)\mathbf{I} + \mathbf{Y}_b^T \mathbf{R}^{-1} \mathbf{Y}_b \right]^{-1} \mathbf{Y}_b^T \mathbf{R}^{-1}$$

- ▶ the deterministic analysis is obtained on the same grid as the ensemble is running on; the *analysis increments* can be interpolated to a higher resolution